# Inclusive $\pi 0$ , $\eta$ , and direct photon production at high transverse momentum in p + p and d + Au collisions at $\sqrt{sNN} = 200$ GeV

(STAR Collaboration) Abelev, B. I.; ...; Planinić, Mirko; ...; Poljak, Nikola; ...; Zoulkarneeva, Y.

Source / Izvornik: Physical Review C - Nuclear Physics, 2010, 81

Journal article, Published version Rad u časopisu, Objavljena verzija rada (izdavačev PDF)

https://doi.org/10.1103/PhysRevC.81.064904

Permanent link / Trajna poveznica: https://urn.nsk.hr/urn:nbn:hr:217:056774

Rights / Prava: In copyright/Zaštićeno autorskim pravom.

Download date / Datum preuzimanja: 2024-12-30



Repository / Repozitorij:

Repository of the Faculty of Science - University of Zagreb





# Inclusive $\pi^0$ , $\eta$ , and direct photon production at high transverse momentum in p+p and $d+\mathrm{Au}$ collisions at $\sqrt{s_{NN}}=200~\mathrm{GeV}$

B. I. Abelev, M. M. Aggarwal, <sup>30</sup> Z. Ahammed, <sup>47</sup> A. V. Alakhverdyants, <sup>17</sup> B. D. Anderson, <sup>18</sup> D. Arkhipkin, <sup>3</sup> G. S. Averichev, <sup>17</sup> J. Balewski, <sup>22</sup> O. Barannikova, <sup>8</sup> L. S. Barnby, <sup>2</sup> S. Baumgart, <sup>52</sup> D. R. Beavis, <sup>3</sup> R. Bellwied, <sup>50</sup> F. Benedosso, <sup>27</sup> M. J. Betancourt, <sup>22</sup> R. R. Betts, <sup>8</sup> A. Bhasin, <sup>16</sup> A. K. Bhati, <sup>30</sup> H. Bichsel, <sup>49</sup> J. Bielcik, <sup>10</sup> J. Bielcikova, <sup>11</sup> B. Biritz, <sup>6</sup> L. C. Bland, <sup>3</sup> B. E. Bonner, <sup>36</sup> J. Bouchet, <sup>18</sup> E. Braidot, <sup>27</sup> A. V. Brandin, <sup>25</sup> A. Bridgeman, <sup>1</sup> E. Bruna, <sup>52</sup> S. Bueltmann, <sup>29</sup> I. Bunzarov, <sup>17</sup> T. P. Burton, X. Z. Cai, H. Caines, M. Calderón de la Barca Sánchez, O. Catu, D. Cebra, R. Cendejas, M. C. Cervantes, <sup>42</sup> Z. Chajecki, <sup>28</sup> P. Chaloupka, <sup>11</sup> S. Chattopadhyay, <sup>47</sup> H. F. Chen, <sup>38</sup> J. H. Chen, <sup>40</sup> J. Y. Chen, <sup>51</sup> J. Cheng, <sup>44</sup> M. Cherney, <sup>9</sup> A. Chikanian, <sup>52</sup> K. E. Choi, <sup>34</sup> W. Christie, <sup>3</sup> P. Chung, <sup>11</sup> R. F. Clarke, <sup>42</sup> M. J. M. Codrington, <sup>42</sup> R. Corliss, <sup>22</sup> J. G. Cramer, <sup>49</sup> H. J. Crawford, <sup>4</sup> D. Das, <sup>5</sup> S. Dash, <sup>13</sup> A. Davila Leyva, <sup>43</sup> L. C. De Silva, <sup>50</sup> R. R. Debbe, <sup>3</sup> T. G. Dedovich, <sup>17</sup> M. DePhillips, <sup>3</sup> A. A. Derevschikov, <sup>32</sup> R. Derradi de Souza, <sup>7</sup> L. Didenko, <sup>3</sup> P. Djawotho, <sup>42</sup> S. M. Dogra, <sup>16</sup> X. Dong, <sup>21</sup> J. L. Drachenberg, <sup>42</sup> J. E. Draper, <sup>5</sup> J. C. Dunlop, <sup>3</sup> M. R. Dutta Mazumdar, <sup>47</sup> L. G. Efimov, <sup>17</sup> E. Elhahluli, <sup>2</sup> M. Elnimr, <sup>50</sup> J. C. Dunlop, <sup>3</sup> M. R. Dutta Mazumdar, <sup>47</sup> L. G. Efimov, <sup>17</sup> E. Elhahluli, <sup>2</sup> M. Elnimr, <sup>50</sup> J. C. Dunlop, <sup>3</sup> M. R. Dutta Mazumdar, <sup>47</sup> L. G. Efimov, <sup>18</sup> E. Elhahluli, <sup>2</sup> M. Elnimr, <sup>50</sup> J. C. Dunlop, <sup>3</sup> M. R. Dutta Mazumdar, <sup>47</sup> L. G. Efimov, <sup>18</sup> E. Elhahluli, <sup>2</sup> M. Elnimr, <sup>50</sup> J. C. Dunlop, <sup>3</sup> M. R. Dutta Mazumdar, <sup>47</sup> L. G. Efimov, <sup>18</sup> E. Elhahluli, <sup>2</sup> M. Elnimr, <sup>50</sup> J. C. Dunlop, <sup>3</sup> M. R. Dutta Mazumdar, <sup>48</sup> D. Elnimr, <sup>50</sup> J. C. Dunlop, <sup>3</sup> M. R. Dutta Mazumdar, <sup>49</sup> L. G. Efimov, <sup>19</sup> E. Elhahluli, <sup>2</sup> M. Elnimr, <sup>50</sup> J. C. Dunlop, <sup>3</sup> M. R. Dutta Mazumdar, <sup>40</sup> L. G. Efimov, <sup>41</sup> D. C. Dunlop, <sup>42</sup> D. C. Dunlop, <sup>43</sup> D. C. Dunlop, <sup>44</sup> D. C. Dunlop, <sup>44</sup> D. C. Dunlop, <sup>44</sup> D. C. Dunlop, <sup>44</sup> D. C. Dunlop, <sup>45</sup> D. C. Dunlop, <sup>45</sup> D. C. Dunlop, <sup>46</sup> D. C. Dunlop, <sup>48</sup> D. C. Dunlop, J. Engelage, <sup>4</sup> G. Eppley, <sup>36</sup> B. Erazmus, <sup>41</sup> M. Estienne, <sup>41</sup> L. Eun, <sup>31</sup> P. Fachini, <sup>3</sup> R. Fatemi, <sup>19</sup> J. Fedorisin, <sup>17</sup> R. G. Fersch, <sup>19</sup> P. Filip, <sup>17</sup> E. Finch, <sup>52</sup> V. Fine, <sup>3</sup> Y. Fisyak, <sup>3</sup> C. A. Gagliardi, <sup>42</sup> D. R. Gangadharan, <sup>6</sup> M. S. Ganti, <sup>47</sup> E. J. Garcia-Solis, <sup>8</sup> A. Geromitsos, <sup>41</sup> F. Geurts, <sup>36</sup> V. Ghazikhanian, <sup>6</sup> P. Ghosh, <sup>47</sup> Y. N. Gorbunov, <sup>9</sup> A. Gordon, <sup>3</sup> O. Grebenyuk, <sup>21</sup> D. Grosnick, <sup>46</sup> B. Grube, <sup>34</sup> S. M. Guertin, <sup>6</sup> A. Gupta, <sup>16</sup> N. Gupta, <sup>16</sup> W. Guryn, <sup>3</sup> B. Haag, <sup>5</sup> T. J. Hallman, <sup>3</sup> A. Hamed, <sup>42</sup> L.-X. Han, <sup>40</sup> J. W. Harris, <sup>52</sup> J. P. Hays-Wehle, <sup>22</sup> M. Heinz, <sup>52</sup> S. Heppelmann, <sup>31</sup> A. Hirsch, <sup>33</sup> E. Hjort, <sup>21</sup> A. M. Hoffman, <sup>22</sup> G. W. Hoffmann, <sup>43</sup> D. J. Hofman, <sup>8</sup> R. S. Hollis, <sup>8</sup> H. Z. Huang, <sup>8</sup> T. J. Humanic, <sup>28</sup> L. Huo, <sup>42</sup> G. Igo, <sup>6</sup> A. Iordanova, <sup>8</sup> P. Jacobs, <sup>21</sup> W. W. Jacobs, <sup>15</sup> P. Jakl, <sup>11</sup> C. Jena, <sup>13</sup> F. Jin, <sup>40</sup> C. L. Jones, <sup>22</sup> P. G. Jones, <sup>2</sup> J. Joseph, <sup>18</sup> E. G. Judd, <sup>4</sup> S. Kabana, <sup>41</sup> K. Kajimoto, <sup>43</sup> K. Kang, <sup>44</sup> J. Kapitan, <sup>11</sup> K. Kauder, <sup>8</sup> D. Keane, <sup>18</sup> A. Kechechyan, <sup>17</sup> D. Kettler, <sup>49</sup> D. P. Kikola, <sup>21</sup> J. Kiryluk, <sup>21</sup> A. Kisiel, <sup>48</sup> S. R. Klein, <sup>21</sup> A. G. Knospe, <sup>52</sup> A. Kocoloski, <sup>22</sup> D. D. Koetke, <sup>46</sup> T. Kollegger, <sup>12</sup> J. Konzer, <sup>33</sup> M. Kopytine, <sup>18</sup> I. Koralt, <sup>29</sup> W. Korsch, <sup>19</sup> L. Kotchenda, <sup>25</sup> V. Kouchpil, <sup>11</sup> P. Kravtsov, <sup>25</sup> K. Krueger, <sup>1</sup> M. Krus, <sup>10</sup> L. Kumar, <sup>30</sup> P. Kurnadi, <sup>6</sup> M. A. C. Lamont, <sup>3</sup> J. M. Landgraf, <sup>3</sup> S. LaPointe, <sup>50</sup> J. Lauret, <sup>3</sup> A. Lebedev, <sup>3</sup> R. Lednicky, <sup>17</sup> C.-H. Lee, <sup>34</sup> J. H. Lee, <sup>3</sup> W. Leight, <sup>22</sup> M. J. LeVine, <sup>3</sup> C. Li <sup>38</sup> I. Li <sup>40</sup> X. Li <sup>33</sup> X. Li <sup>39</sup> Y. Li <sup>44</sup> Z. Li <sup>51</sup> G. Lin, <sup>52</sup> S. J. Lindenbaum, <sup>26,\*</sup> M. A. Lisa, <sup>28</sup> F. Liu, <sup>51</sup> P. Filip, <sup>17</sup> E. Finch, <sup>52</sup> V. Fine, <sup>3</sup> Y. Fisyak, <sup>3</sup> C. A. Gagliardi, <sup>42</sup> D. R. Gangadharan, <sup>6</sup> M. S. Ganti, <sup>47</sup> E. J. Garcia-Solis, <sup>8</sup> C. Li, <sup>38</sup> L. Li, <sup>43</sup> N. Li, <sup>51</sup> W. Li, <sup>40</sup> X. Li, <sup>33</sup> X. Li, <sup>39</sup> Y. Li, <sup>44</sup> Z. Li, <sup>51</sup> G. Lin, <sup>52</sup> S. J. Lindenbaum, <sup>26,\*</sup> M. A. Lisa, <sup>28</sup> F. Liu, <sup>51</sup> C. Li, <sup>36</sup> L. Li, <sup>37</sup> N. Li, <sup>37</sup> W. Li, <sup>36</sup> X. Li, <sup>37</sup> X. Li, <sup>37</sup> Y. Li, <sup>34</sup> Z. Li, <sup>37</sup> G. Lin, <sup>32</sup> S. J. Lindenbaum, <sup>38</sup> M. A. Lisa, <sup>26</sup> F. Liu, <sup>31</sup> H. Liu, <sup>5</sup> J. Liu, <sup>36</sup> T. Ljubicic, <sup>3</sup> W. J. Llope, <sup>36</sup> R. S. Longacre, <sup>3</sup> W. A. Love, <sup>3</sup> Y. Lu, <sup>38</sup> G. L. Ma, <sup>40</sup> Y. G. Ma, <sup>40</sup> D. P. Mahapatra, <sup>13</sup> R. Majka, <sup>52</sup> O. I. Mall, <sup>5</sup> L. K. Mangotra, <sup>16</sup> R. Manweiler, <sup>46</sup> S. Margetis, <sup>18</sup> C. Markert, <sup>43</sup> H. Masui, <sup>21</sup> H. S. Matis, <sup>21</sup> Yu. A. Matulenko, <sup>32</sup> D. McDonald, <sup>36</sup> T. S. McShane, <sup>9</sup> A. Meschanin, <sup>32</sup> R. Milner, <sup>22</sup> N. G. Minaev, <sup>32</sup> S. Mioduszewski, <sup>42</sup> A. Mischke, <sup>27</sup> M. K. Mitrovski, <sup>12</sup> B. Mohanty, <sup>47</sup> M. M. Mondal, <sup>47</sup> D. A. Morozov, <sup>32</sup> M. G. Munhoz, <sup>37</sup> B. K. Nandi, <sup>14</sup> C. Nattrass, <sup>52</sup> T. K. Nayak, <sup>47</sup> J. M. Nelson, <sup>2</sup> P. K. Netrakanti, <sup>33</sup> M. J. Ng, <sup>4</sup> L. V. Nogach, <sup>32</sup> S. B. Nurushev, <sup>32</sup> G. Odyniec, <sup>21</sup> A. Ozoroga, <sup>31</sup> H. Olson, <sup>32</sup> N. Olson, <sup>21</sup> M. Posher, <sup>10</sup> R. S. Posn, <sup>15</sup> S. K. Pal, <sup>47</sup> V. Posn, <sup>44</sup> R. V. A. Ogawa,<sup>3</sup> H. Okada,<sup>3</sup> V. Okorokov,<sup>25</sup> D. Olson,<sup>21</sup> M. Pachr,<sup>10</sup> B. S. Page,<sup>15</sup> S. K. Pal,<sup>47</sup> Y. Pandit,<sup>18</sup> Y. Panebratsev,<sup>17</sup> T. Pawlak,<sup>48</sup> T. Peitzmann,<sup>27</sup> V. Perevoztchikov,<sup>3</sup> C. Perkins,<sup>4</sup> W. Peryt,<sup>48</sup> S. C. Phatak,<sup>13</sup> P. Pile,<sup>3</sup> M. Planinic,<sup>53</sup> M. A. Ploskon, <sup>21</sup> J. Pluta, <sup>48</sup> D. Plyku, <sup>29</sup> N. Poljak, <sup>53</sup> A. M. Poskanzer, <sup>21</sup> B. V. K. S. Potukuchi, <sup>16</sup> C. B. Powell, <sup>21</sup> D. Prindle, <sup>49</sup> C. Pruneau, <sup>50</sup> N. K. Pruthi, <sup>30</sup> P. R. Pujahari, <sup>14</sup> J. Putschke, <sup>52</sup> R. Raniwala, <sup>35</sup> S. Raniwala, <sup>35</sup> R. L. Ray, <sup>43</sup> R. Redwine, <sup>22</sup> R. Reed, <sup>5</sup> J. M. Rehberg, <sup>12</sup> H. G. Ritter, <sup>21</sup> J. B. Roberts, <sup>36</sup> O. V. Rogachevskiy, <sup>17</sup> J. L. Romero, <sup>5</sup> A. Rose, <sup>21</sup> C. Roy, <sup>41</sup> L. Ruan, <sup>3</sup> M. J. Russcher, <sup>27</sup> R. Sahoo, <sup>41</sup> S. Sakai, <sup>6</sup> I. Sakrejda, <sup>21</sup> T. Sakuma, <sup>22</sup> S. Salur, <sup>5</sup> J. Sandweiss, <sup>52</sup> E. Sangaline, <sup>5</sup> J. Schambach, <sup>43</sup> R. P. Scharenberg, <sup>33</sup> N. Schmitz, <sup>23</sup> T. R. Schuster, <sup>12</sup> J. Seele, <sup>22</sup> J. Seger, <sup>9</sup> I. Selyuzhenkov, <sup>15</sup> P. Seyboth, <sup>23</sup> E. Shahaliev, <sup>17</sup> M. Gland, <sup>38</sup> M. G. Sakai, <sup>50</sup> G. G. G. G. Sakai, <sup>51</sup> F. P. Girler, <sup>23</sup> R. N. Schuster, <sup>23</sup> R. N. Schuster, <sup>23</sup> R. N. Schuster, <sup>23</sup> R. N. Schuster, <sup>24</sup> R. G. Sakai, <sup>52</sup> R. Sakai, <sup>53</sup> R. Sakai, <sup>54</sup> R. Sakai, <sup>54</sup> R. Sakai, <sup>55</sup> R. Sakai, <sup>55</sup> R. Schuster, <sup>26</sup> R. Sakai, <sup>56</sup> R. Sakai, <sup>57</sup> R. Schuster, <sup>27</sup> R. Schuster, <sup>28</sup> R. Sakai, <sup>58</sup> R. Sakai, <sup>58</sup> R. Sakai, <sup>58</sup> R. Schuster, <sup>28</sup> R. Sakai, <sup>58</sup> R. Sakai, <sup>59</sup> R. Sakai, <sup>59</sup> R. Sakai, <sup>59</sup> R. Sakai, <sup>59</sup> R. Sakai, <sup>50</sup> R. Saka M. Shao, <sup>38</sup> M. Sharma, <sup>50</sup> S. S. Shi, <sup>51</sup> E. P. Sichtermann, <sup>21</sup> F. Simon, <sup>23</sup> R. N. Singaraju, <sup>47</sup> M. J. Skoby, <sup>33</sup> N. Smirnov, <sup>52</sup> P. Sorensen, <sup>3</sup> J. Sowinski, <sup>15</sup> H. M. Spinka, <sup>1</sup> B. Srivastava, <sup>33</sup> T. D. S. Stanislaus, <sup>46</sup> D. Staszak, <sup>6</sup> J. R. Stevens, <sup>15</sup> R. Stock, <sup>12</sup> M. Strikhanov, <sup>25</sup> B. Stringfellow, <sup>33</sup> A. A. P. Suaide, <sup>37</sup> M. C. Suarez, <sup>8</sup> N. L. Subba, <sup>18</sup> M. Sumbera, <sup>11</sup> X. M. Sun, <sup>21</sup> Y. Sun, <sup>38</sup> Z. Sun, <sup>20</sup> B. Surrow, <sup>22</sup> T. J. M. Symons, <sup>21</sup> A. Szanto de Toledo, <sup>37</sup> J. Takahashi, <sup>7</sup> A. H. Tang, <sup>3</sup> Z. Tang, <sup>38</sup> L. H. Tarini, <sup>50</sup> T. Tang, <sup>50</sup> T. Tarnowsky, <sup>24</sup> D. Thein, <sup>43</sup> J. H. Thomas, <sup>21</sup> J. Tian, <sup>40</sup> A. R. Timmins, <sup>50</sup> S. Timoshenko, <sup>25</sup> D. Tlusty, <sup>11</sup> M. Tokarev, <sup>17</sup> T. A. Trainor, <sup>49</sup> V. N. Tram, <sup>21</sup> S. Trentalange, <sup>6</sup> R. E. Tribble, <sup>42</sup> O. D. Tsai, <sup>6</sup> J. Ulery, <sup>33</sup> T. Ullrich, <sup>3</sup> D. G. Underwood, <sup>1</sup> G. Van Buren,<sup>3</sup> G. van Nieuwenhuizen,<sup>22</sup> J. A. Vanfossen Jr.,<sup>18</sup> R. Varma,<sup>14</sup> G. M. S. Vasconcelos,<sup>7</sup> A. N. Vasiliev,<sup>32</sup> F. Videbaek,<sup>3</sup> Y. P. Viyogi,<sup>47</sup> S. Vokal,<sup>17</sup> S. A. Voloshin,<sup>50</sup> M. Wada,<sup>43</sup> M. Walker,<sup>22</sup> F. Wang,<sup>33</sup> G. Wang,<sup>6</sup> H. Wang,<sup>24</sup> J. S. Wang,<sup>20</sup> Q. Wang,<sup>33</sup> X. Wang,<sup>44</sup> X. L. Wang,<sup>38</sup> Y. Wang,<sup>44</sup> G. Webb,<sup>19</sup> J. C. Webb,<sup>46</sup> G. D. Westfall,<sup>24</sup> C. Whitten Jr.,<sup>6</sup> H. Wieman,<sup>21</sup> E. Wingfield,<sup>43</sup> S. W. Wissink,<sup>15</sup> R. Witt,<sup>45</sup> Y. Wu,<sup>51</sup> W. Xie,<sup>33</sup> N. Xu,<sup>21</sup> Q. H. Xu,<sup>39</sup> W. Xu,<sup>6</sup> Y. Xu,<sup>38</sup> Z. Xu,<sup>3</sup> L. Xue,<sup>40</sup> Y. Yang,<sup>20</sup> P. Yepes,<sup>36</sup> K. Yip,<sup>3</sup> I.-K. Yoo,<sup>34</sup> Q. Yue,<sup>44</sup> M. Zawisza,<sup>48</sup> H. Zbroszczyk,<sup>48</sup> W. Zhan,<sup>20</sup> S. Zhang,<sup>40</sup> W. M. Zhang,<sup>18</sup> X. P. Zhang,<sup>21</sup> Z. P. Zhang,<sup>38</sup> J. Zhao,<sup>40</sup> C. Zhong,<sup>40</sup> J. Zhou,<sup>36</sup> W. Zhou,<sup>39</sup> X. Zhu,<sup>44</sup> Y. H. Zhu,<sup>40</sup> R. Zoulkarneev, <sup>17</sup> and Y. Zoulkarneeva <sup>17</sup>

# (STAR Collaboration)

<sup>1</sup>Argonne National Laboratory, Argonne, Illinois 60439, USA

<sup>2</sup>University of Birmingham, Birmingham, United Kingdom

<sup>3</sup>Brookhaven National Laboratory, Upton, New York 11973, USA

<sup>4</sup>University of California, Berkeley, California 94720, USA

<sup>5</sup>University of California, Davis, California 95616, USA

<sup>6</sup>University of California, Los Angeles, California 90095, USA

<sup>7</sup>Universidade Estadual de Campinas, Sao Paulo, Brazil

```
<sup>8</sup>University of Illinois at Chicago, Chicago, Illinois 60607, USA
                  <sup>9</sup>Creighton University, Omaha, Nebraska 68178, USA
<sup>10</sup>Czech Technical University in Prague, FNSPE, Prague CZ-115 19, Czech Republic
        <sup>11</sup>Nuclear Physics Institute AS CR, 250 68 Řež/Prague, Czech Republic
                      <sup>12</sup>University of Frankfurt, Frankfurt, Germany
                    <sup>13</sup>Institute of Physics, Bhubaneswar 751005, India
                     <sup>14</sup>Indian Institute of Technology, Mumbai, India
                <sup>15</sup>Indiana University, Bloomington, Indiana 47408, USA
                       <sup>16</sup>University of Jammu, Jammu 180001, India
           <sup>17</sup>Joint Institute for Nuclear Research, Dubna RU-141 980, Russia
                     <sup>18</sup>Kent State University, Kent, Ohio 44242, USA
            <sup>19</sup>University of Kentucky, Lexington, Kentucky, 40506-0055, USA
                      <sup>20</sup>Institute of Modern Physics, Lanzhou, China
     <sup>21</sup>Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA
<sup>22</sup>Massachusetts Institute of Technology, Cambridge, Massachusetts 02139-4307, USA
                   <sup>23</sup>Max-Planck-Institut für Physik, Munich, Germany
           <sup>24</sup>Michigan State University, East Lansing, Michigan 48824, USA
                <sup>25</sup>Moscow Engineering Physics Institute, Moscow, Russia
           <sup>26</sup>City College of New York, New York City, New York 10031, USA
            <sup>27</sup>NIKHEF and Utrecht University, Amsterdam, The Netherlands
                  <sup>28</sup>Ohio State University, Columbus, Ohio 43210, USA
                <sup>29</sup>Old Dominion University, Norfolk, Virginia 23529, USA
                     <sup>30</sup>Panjab University, Chandigarh 160014, India
     <sup>31</sup>Pennsylvania State University, University Park, Pennsylvania 16802, USA
                   <sup>32</sup>Institute of High Energy Physics, Protvino, Russia
                <sup>33</sup>Purdue University, West Lafayette, Indiana 47907, USA
                 <sup>34</sup>Pusan National University, Pusan, Republic of Korea
                     <sup>35</sup>University of Rajasthan, Jaipur 302004, India
                      <sup>36</sup>Rice University, Houston, Texas 77251, USA
                     <sup>37</sup>Universidade de Sao Paulo, Sao Paulo, Brazil
          <sup>38</sup>University of Science & Technology of China, Hefei 230026, China
                 <sup>39</sup>Shandong University, Jinan, Shandong 250100, China
           <sup>40</sup>Shanghai Institute of Applied Physics, Shanghai 201800, China
                               <sup>41</sup>SUBATECH, Nantes, France
              <sup>42</sup>Texas A&M University, College Station, Texas 77843, USA
                     <sup>43</sup>University of Texas, Austin, Texas 78712, USA
                      <sup>44</sup>Tsinghua University, Beijing 100084, China
           <sup>45</sup>United States Naval Academy, Annapolis, Maryland 21402, USA
                <sup>46</sup>Valparaiso University, Valparaiso, Indiana 46383, USA
               <sup>47</sup>Variable Energy Cyclotron Centre, Kolkata 700064, India
                   <sup>48</sup>Warsaw University of Technology, Warsaw, Poland
              <sup>49</sup>University of Washington, Seattle, Washington 98195, USA
                <sup>50</sup>Wayne State University, Detroit, Michigan 48201, USA
         <sup>51</sup>Institute of Particle Physics, CCNU (HZNU), Wuhan 430079, China
                 <sup>52</sup>Yale University, New Haven, Connecticut 06520, USA
                    <sup>53</sup>University of Zagreb, Zagreb HR-10002, Croatia
                 (Received 18 December 2009; published 15 June 2010)
```

We report a measurement of high- $p_T$  inclusive  $\pi^0$ ,  $\eta$ , and direct photon production in p+p and d+Au collisions at  $\sqrt{s_{NN}}=200$  GeV at midrapidity ( $0<\eta<1$ ). Photons from the decay  $\pi^0\to\gamma\gamma$  were detected in the barrel electromagnetic calorimeter of the STAR experiment at the Relativistic Heavy Ion Collider. The  $\eta\to\gamma\gamma$  decay was also observed and constituted the first  $\eta$  measurement by STAR. The first direct photon cross-section measurement by STAR is also presented; the signal was extracted statistically by subtracting the  $\pi^0$ ,  $\eta$ , and  $\omega(782)$  decay background from the inclusive photon distribution observed in the calorimeter. The analysis is described in detail, and the results are found to be in good agreement with earlier measurements and with next-to-leading-order perturbative QCD calculations.

DOI: 10.1103/PhysRevC.81.064904 PACS number(s): 13.85.Ni, 13.85.Qk, 13.87.Fh, 25.75.-q

## I. INTRODUCTION

The high center-of-mass energy ( $\sqrt{s_{NN}} = 200 \text{ GeV}$ ) of the Relativistic Heavy Ion Collider (RHIC) opens up the hard scattering regime, which is accessed by measuring particle production at high transverse momentum  $p_T$ . The high- $p_T$ particles ( $p_T \gtrsim 3 \text{ GeV}/c$ ) originate from the fragmentation of partons that have scattered in the early stage of the collisions. Hence, in heavy-ion collisions the high- $p_T$  particles can be used to probe the produced medium of strongly interacting matter. A significant suppression of high- $p_T$  hadron production relative to a simple binary collision scaling from p + phas been observed at RHIC in central Au + Au collisions [1]. Furthermore, it was found that jetlike correlations opposite to trigger jets are suppressed and that the azimuthal anisotropy in hadron emission persists out to very high  $p_T$  [2-4]. In contrast, no suppression effects were seen in d + Au collisions [5–8], which has led to the conclusion that the observations made in Au + Au are attributable to the high-density medium produced in such collisions and not to initial-state effects. The most probable explanation to date is that the suppression is attributable to parton energy loss from induced gluon radiation (jet quenching) in the extremely hot and dense medium [9]. To quantitatively understand this behavior and, in particular, to separate hot from cold nuclear-matter effects, such as the Cronin effect [10] and parton shadowing and antishadowing [11–13], precise measurements of identified hadrons at high  $p_T$  in p + p and d + Au collisions are required [14].

Prompt photons have long been proposed as a powerful tool for studying the jet quenching via photon-jet correlations [15]. In the dominant hard photon production processes (quark-gluon Compton scattering and quark-antiquark annihilation), the outgoing photon balances the momentum of its partner parton and has large enough mean free path to escape the collision system, providing a calibrated probe for studying the energy loss and mean free path of the parton in the medium. In addition, prompt photons constitute a background for measuring the medium-induced production of photons in response to the energy deposited by that parton [16].

The thermal photon spectrum is directly related to the temperature of the hot and dense medium created in the heavyion collision, provided that it is in thermal equilibrium [17]. The measurement of such a spectrum requires a knowledge of the prompt photon background, which can be measured in p+p and  $d+\mathrm{Au}$  systems that share multiple sources of photons with heavy-ion collisions but are not expected to produce an extended thermal system.

The measurements of  $\pi^0$ 's and direct photons in p+p collisions are also of specific interest for studies of the proton spin structure (see, e.g., Ref. [18]), which are under way at RHIC. A main objective of the RHIC spin program is to constrain the polarization of the gluons inside the proton  $[\Delta G(x)]$ . The unpolarized cross sections provide a test of the next-to-leading-order perturbative QCD (NLO pQCD) framework, which is used to interpret the measured spin-dependent observables.

In this article, we present the first results for the high- $p_T \pi^0$ ,  $\eta$ , and direct photon production in p + p and d + Au collisions at  $\sqrt{s_{NN}} = 200$  GeV in the pseudorapidity range  $0 < \eta < 1$ , measured by the STAR experiment at RHIC (except the cross section for  $\pi^0$  production in p + p collisions, first presented in Ref. [19]). The STAR barrel electromagnetic calorimeter was used to detect high- $p_T$   $\pi^0$  and  $\eta$  mesons via their  $\gamma \gamma$  decays. The direct photon signal was extracted statistically by subtracting the  $\pi^0$ ,  $\eta$ , and  $\omega(782)$  decay background from the inclusive photon distribution observed in the calorimeter. The presented data constitute a necessary baseline for the measurements of  $\pi^0$ ,  $\eta$ , and direct photon production in heavy-ion collisions at RHIC. Inclusive  $\pi^0$ production was previously measured in STAR for low  $p_T$  at midrapidity in Au + Au collisions at  $\sqrt{s_{NN}} = 130 \text{ GeV}$  [20] and 200 GeV [21], and at the forward rapidities in p + p and  $d + \text{Au collisions at } \sqrt{s_{NN}} = 200 \text{ GeV [22]}$ . STAR has also measured the production of other identified particles, such as  $\pi^{\pm}$ ,  $K^{\pm}$ ,  $p/\bar{p}$ , and hadronic resonances [14,23–25]. The PHENIX experiment at RHIC has also measured the  $\pi^0$ ,  $\eta$ , and direct photon production at  $\sqrt{s_{NN}} = 200 \text{ GeV}$  in a variety

The article is organized as follows. In Sec. II, we describe the detectors that were used in this analysis. In Sec. III, we describe the data-processing chain used to reconstruct photon candidates in the raw data. Sections IV and V show how these photon candidates were used to calculate the yields of  $\pi^0$  and  $\eta$  and direct photons, respectively. Finally, in Sec. VI, we present the results and compare our data to the theoretical calculations and to the measurements by other experiments.

of collision systems, including p + p and d + Au [6,26-29].

# II. EXPERIMENTAL SETUP

The STAR detector (solenoidal tracker at RHIC) [30] was designed primarily for measurements of hadron production in heavy-ion and proton-proton collisions over a large solid angle. For this purpose, tracking detectors with large acceptance and high granularity were placed inside a large-volume solenoidal magnetic field (0.5 T). The detector subsystems relevant for the present analysis are briefly described in the following sections.

# A. Time projection chamber

The time projection chamber (TPC) [31] is the central tracking device in STAR. It allows one to track charged particles, measure their momenta, and identify the particle species by measuring the ionization energy loss dE/dx.

The TPC barrel measures 4.2 m in length and has an inner radius of 0.5 m and an outer radius of 2 m. The TPC acceptance covers  $\pm 1.8$  units in pseudorapidity and full azimuth. Particle momentum is measured in the range 0.1–30 GeV/c. In this analysis, TPC tracks were used to reconstruct the interaction vertex and to identify the energy deposits of charged particles in the calorimeter.

# **B.** Forward TPC modules

Two forward time projection chambers (FTPCs) [32] extend the STAR tracking capability to the pseudorapidity range  $2.5 < |\eta| < 4$ . Each FTPC is a cylindrical volume with a

<sup>\*</sup>Deceased.

diameter of 75 cm and a length of 120 cm, with radial drift field and pad readout chambers mounted on the outer cylindrical surface. Two such detectors were installed partially inside the main TPC, on both sides of the interaction point. In this analysis, the forward charged-track multiplicity recorded in the FTPC in the gold beam direction served as a measure of the centrality in d+ Au collisions.

#### C. Barrel electromagnetic calorimeter

A barrel electromagnetic calorimeter (BEMC) [33] was incrementally added to the STAR setup in 2001–2005 to measure the energy deposited by high- $p_T$  photons and electrons and to provide a trigger signal. The calorimeter is located inside the magnet coil and surrounds the TPC, covering a pseudorapidity range  $|\eta| < 1$  and full azimuth.

The full calorimeter consists of two contiguous half barrels, located east and west of the nominal interaction point, each of which is azimuthally segmented into 60 modules. Each module is approximately 26 cm wide and covers  $6^{\circ}$  (105 mrad) in azimuth and one unit in pseudorapidity. The active depth is 23.5 cm, to which 6.6 cm of structural elements are added at the outer radius. Results presented in this article used only the west calorimeter half barrel (0 <  $\eta$  < 1), which was fully installed and calibrated in 2003–2005.

The modules are segmented into 40 projective towers of lead-scintillator stacks, 2 in the  $\varphi$  and 20 in the  $\eta$  direction. A tower covers 0.05 rad in  $\Delta \varphi$  and 0.05 units in  $\Delta \eta$ . Each calorimeter half barrel is thus segmented into a total of 2400 towers. Each tower consists of a stack of 20 layers of lead and 21 layers of scintillator. All these layers are 5 mm thick, except the first two scintillator layers, which are 6 mm thick. A separate readout of these two layers provides the calorimeter preshower signal, which was not used in this analysis. A shower maximum detector (see the following section) is positioned behind the fifth scintillator layer. The whole stack is held together by mechanical compression and friction between layers. From layer-by-layer tests of the BEMC optical system, together with an analysis of cosmic ray and beam test data, the nominal energy resolution of the calorimeter is estimated to be  $\delta E/E = 14\%/\sqrt{E(\text{GeV})} \oplus 1.5\%$  [33].

#### D. Shower maximum detector

The shower maximum detector (SMD) is a multiwire proportional counter with strip readout. It is located at a depth of approximately 5.6 radiation lengths at  $\eta=0$ , increasing to 7.9 radiation lengths at  $\eta=1$ , including all material immediately in front of the calorimeter. The purpose of the SMD is to improve the spatial resolution of the calorimeter and to measure the shower profile. This is necessary because the transverse dimension of each tower  $(\approx\!10\times10~{\rm cm}^2)$  is much larger than the lateral spread of an electromagnetic shower. The improved resolution is essential for separating the two photon showers originating from the decay of high-momentum  $\pi^0$  and  $\eta$  mesons.

Independent cathode planes with strips along the  $\eta$  and  $\varphi$  directions allow the reconstruction of two projections of a

shower. The coverage is  $\Delta \eta \times \Delta \varphi = 0.0064 \times 0.1$  rad for the  $\eta$  strips and  $0.1 \times 0.0064$  rad for the  $\varphi$  strips, with each group of  $2 \times 2$  towers covering 15 strips in each SMD plane behind it. In total, the SMD contains 36 000 strips.

## E. Trigger detectors

In addition to the STAR barrel detectors, sampling hadronic calorimeters were placed at a distance of 18 m from the interaction point on both sides of the experimental hall. In heavy-ion collisions, these zero-degree calorimeters (ZDCs) [34,35] measure the total energy of the unbound neutrons emitted from the nuclear fragments after a collision. The charged fragments of the collision are bent away by the RHIC dipole magnets upstream of the ZDCs. For the  $d+\mathrm{Au}$  data used in this analysis, the ZDC provided a collision trigger by requiring the detection of at least one neutron in the gold beam direction.

To provide a collision trigger in p+p collisions, beambeam counters (BBCs) [36,37] were mounted around the beam pipe beyond both pole tips of the STAR magnet at a distance of 3.7 m from the interaction point. The detector consists of two sets of small and large hexagonal scintillator tiles arranged into a ring that covers pseudorapidities between 2.1 and 5.0. The minimum bias trigger required a coincidence of signals in at least one of the 18 small BBC tiles on each side of the interaction region.

The two BBCs record timing signals that can be used to determine the time of flight for the forward fragments. The difference between these two flight times provides a measurement of the z position of the interaction vertex ( $z_{\rm vert}$ ) to an accuracy of about 40 cm [38]. Events with large values of the time-of-flight difference, which indicate the passage of beam background, were rejected at the trigger level. The BBCs also served to measure the beam luminosity in p+p runs.

# III. DATA RECONSTRUCTION

#### A. Datasets and statistics

The data used in this analysis were taken in the  $d+\mathrm{Au}$  run of 2003 and in the p+p run of 2005, both at  $\sqrt{s_{NN}}=200$  GeV. The integrated luminosity was 0.66 pb<sup>-1</sup> for the p+p data and the equivalent nucleon-nucleon luminosity was 0.22 pb<sup>-1</sup> for the  $d+\mathrm{Au}$  data. The following trigger conditions were used.

# 1. Minimum bias (MinBias) trigger in d + Au collisions

This condition required the presence of at least one neutron signal in the ZDC in the gold beam direction. As determined from detailed simulations of the ZDC acceptance [5], this trigger captured (95  $\pm$  3)% of the total d + Au hadronic cross section of  $\sigma_{\rm hadr}^{d+{\rm Au}}=2.21\pm0.09{\rm b}$ .

# 2. MinBias trigger in p + p collisions

This condition required the coincidence of signals from two BBC tiles on opposing sides of the interaction point. Due to the dual-arm configuration, this trigger was sensitive to the nonsingly diffractive (NSD) cross section, which is a sum of the nondiffractive and doubly diffractive cross sections. The total inelastic cross section is a sum of the NSD and singly diffractive cross sections. A MinBias cross section of  $\sigma_{\rm BBC}=26.1\pm0.2\,({\rm stat.})\pm1.8\,({\rm syst.})$  mb was independently measured via van der Meer scans in dedicated accelerator runs [39]. This trigger captured (87  $\pm$  8)% of the p+p NSD cross section, as was determined from a detailed simulation of the BBC acceptance [1]. Correcting the BBC cross section for the acceptance, we obtained the NSD cross section  $\sigma_{\rm NSD}^{p+p}=30.0\pm3.5\,{\rm mb}.$ 

## 3. HighTower trigger

This condition required a transverse energy deposit  $E_T$  above a predefined threshold in at least one calorimeter tower, in addition to satisfying the MinBias condition. This trigger enriched the recorded dataset with events that had a large  $E_T$ . Two different thresholds were applied, defining the HighTower-1 and HighTower-2 datasets. The nominal values of these thresholds were set to 2.6 and 3.5 GeV in p+p and to 2.5 and 4.5 GeV in d+Au runs. Prior to each run, all towers were equalized to give the uniform transverse energy response by adjusting the high-voltage settings of the individual photomultipliers.

# 4. HighTower software filter

The HighTower-triggered data were additionally filtered using a software implementation of the HighTower trigger. In this filter, the highest tower ADC value found in the event was required to exceed the same HighTower-1 (HighTower-2) threshold as the one that was used during the run. This filter was needed to remove events that were falsely triggered due to the presence of noisy channels (hot towers). Such channels were identified off line in a separate analysis and recorded in a database. In addition, the highest calibrated transverse energy of a tower in the event was required to exceed slightly higher thresholds ( $E_T + 0.5 \text{ GeV}$ ) than those used during the run, to account for possible inaccuracy of the online calibration of the towers. This software filter also served to make the trigger efficiency for the Monte Carlo detector simulation and for the real data as close as possible.

# B. Beam background rejection

During the data taking in 2003–2005, interactions of beam ions with material approximately 40 m upstream from the interaction region gave rise to particles that traversed the detector almost parallel to the beam direction. This source of background was eliminated by installing additional shielding in the RHIC tunnel for the subsequent runs.

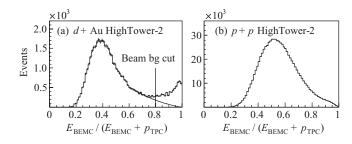


FIG. 1. (a) Distribution of  $r = E_{\rm BEMC}/(E_{\rm BEMC} + p_{\rm TPC})$  in  $d+{\rm Au}$  events, which shows beam background at r>0.8. The curve corresponds to a second-order polynomial fit, constrained to pass through zero at r=1, used to estimate the false rejection rate. (b) Distribution of r in p+p events.

To identify events containing such background tracks, the ratio

$$r = \frac{E_{\text{BEMC}}}{E_{\text{BEMC}} + p_{\text{TPC}}} \tag{1}$$

was calculated, where  $E_{\text{BEMC}}$  is the total transverse energy recorded in the BEMC and  $p_{TPC}$  is the transverse momentum sum of all charged tracks reconstructed in the TPC. In events containing background, r was large (close to 1) because photons from these background events deposited a large amount of energy in the calorimeter, while the accompanying charged tracks were not reconstructed in the TPC, because they did not point to the vertex. Figure 1 shows the distributions of r for the d + Au and p + p data. The peak near unity in panel (a) indicates the presence of beam background in d + Au collisions. Events with r > 0.8 were removed from the d + Au analysis. This cut rejected 3.4% of MinBias and 13% of HighTower-2 events. From a polynomial fit to the d + Audistribution in the region r = 0.6–0.8 [curve in Fig. 1(a)], the false rejection rate was estimated to be 3.6% in the d + AuHighTower-2 data and less than 1% in the other datasets. By studying this rejection rate as a function of  $E_{\text{BEMC}}$ , we estimated the potential distortion of the  $\pi^0$ ,  $\eta$ , and photon spectra due to the removal of these events to be below 1% in all datasets.

Figure 1(b) shows the distribution of r for the HighTower p+p data. The background was negligible because of the BBC coincidence requirement in the trigger and the timing cut on the BBC vertex position. Therefore, no cut on r was applied to the p+p data.

The residual beam background contamination in the  $d+\mathrm{Au}$  MinBias trigger was estimated from an analysis of the empty RHIC bunches to be  $(5\pm1)\%$  [14]. To estimate the residual background in our data, we analyzed a sample of  $3\times10^5$  MinBias triggers from unpaired RHIC bunches. These events were passed through the same reconstruction procedure as other data. We observed that  $\approx10\%$  of the fake triggers passed all cuts and that none of these contained a reconstructed  $\pi^0$ . The residual beam background contamination in the  $\pi^0$  yield was thus estimated to be  $0.1\times5\%=0.5\%$  and considered to be negligible.

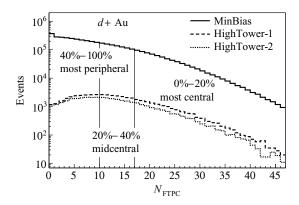


FIG. 2. Centrality selection in the d+Au data based on the FTPC multiplicity  $N_{\rm FTPC}$ . Three centrality classes were defined, containing 0%–20% most central, 20%–40% midcentral, and 40%–100% most peripheral events, respectively.

#### C. Determination of centralities

To measure the centrality in d+Au collisions, we used the correlation between the impact parameter of the collision and the charged-track multiplicity in the forward direction. This correlation was established from a Monte Carlo Glauber simulation [40–42] using, as an input, the Woods-Saxon nuclear-matter density for the gold ion [43] and the Hulthén wave function of the deuteron [44]. In this simulation, the inelastic cross section for a nucleon-nucleon collision was taken to be  $\sigma_{\rm inel}^{NN}=42$  mb. The produced particles were propagated through a full GEANT [45] simulation of the STAR detector. Both the charged-track multiplicity and the number of nucleon-nucleon collisions simulated by the event generator were recorded.

For the event-by-event centrality determination, we measured the multiplicity of tracks reconstructed in the FTPC module in the gold beam direction ( $N_{\rm FTPC}$ ). Centrality bins were defined following the scheme used in other STAR publications [5]. The following quality cuts were applied to the reconstructed tracks: (i) at least six hits were required on the track; (ii)  $p_T < 3~{\rm GeV}/c$ , which guaranteed that the track was fully contained in the FTPC acceptance; and (iii) distance of closest approach to the vertex had to be less than 3 cm. The multiplicity distributions obtained from the  $d+{\rm Au}$  data are shown in Fig. 2 for the MinBias, HighTower-1, and HighTower-2 triggers.

Based on  $N_{\rm FTPC}$ , the events were separated into three centrality classes: 0%–20% most central, 20%–40% midcentral, and 40%–100% most peripheral, as indicated by the vertical lines in Fig. 2. Table I lists the  $N_{\rm FTPC}$  ranges and the corresponding mean numbers of binary collisions ( $\langle N_{\rm coll} \rangle$ ) obtained from the Glauber model for each centrality class. The systematic uncertainties on  $\langle N_{\rm coll} \rangle$  were estimated by varying the Glauber model parameters.

# D. Vertex finding efficiency

In p + p data, a vertex was reconstructed based on the tracking information for 65% of the MinBias events. For the

TABLE I. Centrality classes defined for the d+Au data and the corresponding  $\langle N_{\rm coll} \rangle$  values [5]. The errors given for  $\langle N_{\rm coll} \rangle$  indicate the systematic uncertainty.

Centrality class	N <sub>FTPC</sub> range	$\langle N_{ m coll}  angle$
d + Au MinBias	_	$7.5 \pm 0.4$
0%–20% Most central	≥17	$15.0 \pm 1.1$
20%-40% Midcentral	10-16	$10.2 \pm 1.0$
40%–100% Most peripheral	<10	$4.0 \pm 0.3$
p + p	_	1

remaining events, the vertex position in z was determined using the time information from the BBCs.

In the  $d+\mathrm{Au}$  HighTower data, the charged-track multiplicities were large enough to always have a reconstructed vertex. However, a vertex was missing in about 7% of the MinBias events and cannot be recovered from BBC information because the BBC was not included in the  $d+\mathrm{Au}$  MinBias trigger. Events without a vertex have low charged-track multiplicity, and the contribution from these events to the  $\pi^0$  yield above 1 GeV was assumed to be negligible [46]. Therefore, a correction for vertex inefficiency was applied as a constant normalization factor to the yield and its uncertainty contributed to the total normalization uncertainty of the measured cross sections.

The vertex reconstruction efficiency in triggered  $d+{\rm Au}$  MinBias events was  $\varepsilon_{\rm vert}=0.93\pm0.01$  [5]. However, this efficiency depends on the collision centrality, and we assumed that it was 100% for central events. Scaling the efficiency in the preceding by the ratio of peripheral to total number of  $d+{\rm Au}$  events, we obtained an efficiency correction factor of  $0.88\pm0.02$  for the sample of peripheral events.

Events with  $|z_{\text{vert}}| > 60$  cm were rejected in the analysis because the amount of material traversed by a particle increases dramatically at large values of  $|z_{\text{vert}}|$ . As a consequence, the TPC tracking efficiency is reduced for vertices located far from the center of the detector.

# E. Energy calibration of the calorimeter

In the first step of the calorimeter calibration, the gains of the individual towers were matched to achieve an overall uniform response of the detector. For this purpose, minimum ionizing particles (MIPs) were used, by selecting the TPC tracks of sufficiently large momentum (greater than  $\approx 1~{\rm GeV}/c$ ). These tracks were extrapolated to the BEMC and the response spectra were accumulated, provided that the track extrapolation was contained within one tower and that the track was isolated. For  $d+{\rm Au}$  data, the isolation criterion meant that no other tracks were found in a  $3\times 3$  patch around the tower; for p+p data, these neighboring towers were required to have no signal above noise. The peak positions of such MIP signals were used to calculate the tower-by-tower gain corrections needed to equalize the detector response [47].

In the second step, the energy scale was determined by comparing the momenta p of identified electrons in the TPC with the energies E recorded in the BEMC using the relation

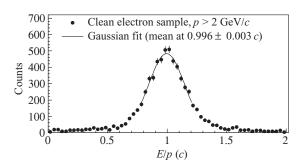


FIG. 3. Electron energy measured in the BEMC after calibration, divided by the momentum measured in the TPC, in the p+p data. The solid line is a Gaussian fit, which shows that the peak is centered at unity.

E=p for ultrarelativistic electrons. Figure 3 shows the distribution of E/p for a selected sample of at least 90% pure electrons in the p+p data at p>2 GeV/c, after the calibration has been performed. The Gaussian fit to the central part of the electron peak demonstrates that the mean has been placed at unity. From a variation of the peak position with p, the systematic uncertainty of the electron calibration was conservatively estimated to be 5%. Within the present statistics, that calibration covers the momentum range only up to p=6 GeV/c. Because the peak position is close to unity at p>3.5 GeV/c, we assume that the assigned systematic uncertainty covers possible nonlinearities at higher photon momenta  $p\lesssim 15$  GeV/c probed in the present measurements.

This calibration method takes advantage of the well understood TPC detector for the precise measurement of the electron track momentum in a wide range. A disadvantage is that it takes large statistics to calibrate the high-energy part of the spectrum. For this reason, only one global calibration constant was obtained. It was found that the current calibration is less reliable at the edges of the half barrel. Therefore, the signals from the two  $\eta$  rings at each side were removed from the analysis.

The absolute energy calibration of the SMD was determined using the beam test data to an accuracy of about 20%. This analysis is not very sensitive to the absolute energy scale of the SMD, because the main energy mesurement was done with the towers.

# F. Particle reconstruction in the BEMC

The first step in the photon reconstruction was to find clusters of energy deposits in the calorimeter by grouping adjacent hits that were likely to have originated from a single incident particle. The cluster-finding algorithm was applied to the signals from BEMC towers and from each of the two SMD layers.

The clustering started from the most energetic hit (seed) in a module and added neighboring hits of decreasing energy to the cluster, until either a predefined maximal cluster size or a lower hit energy threshold was reached. The algorithm then proceeded to process the next seed. The threshold values are listed in Table II. By construction, the clusters were confined within a module and could not be shared by adjacent modules.

TABLE II. Cluster finder threshold values used in the analysis.  $E_{\rm seed}$ ,  $E_{\rm add}$ , and  $E_{\rm min}$  are the minimal energies for the seed hits, regular hits, and entire clusters, respectively.  $N_{\rm max}$  is the maximal cluster size.

Detector	E <sub>seed</sub> (GeV)	E <sub>add</sub> (GeV)	$E_{\min}(\text{GeV})$	$N_{\rm max}$
Towers	0.35	0.035	0.02	4
SMD	0.2	0.0005	0.1	5

However, the likelihood of shower sharing between modules is considered to be low, because the modules are physically separated by  $\approx$ 12-mm-wide air gaps. The  $\eta$ - $\varphi$  position of each cluster was calculated as the energy-weighted mean of the individual hit positions within the cluster.

After the tower and SMD clusters were found, they were combined into BEMC points, which closely corresponded to the impact points and energy deposits of particles that traversed the calorimeter. The procedure for forming the BEMC points is described in detail in Ref. [38]. The SMD information was essential because the minimal opening angle of the decay photons decreases with increasing energy of the parent  $\pi^0$ . The spatial resolution of the BEMC towers alone is not sufficient to efficiently resolve the decay photons of  $\pi^0$ 's with p > 5 GeV/c. For this reason, only the BEMC points that contained tower, SMD- $\eta$ , and SMD- $\varphi$  clusters were kept for the further analysis of the HighTower data. In the analysis of MinBias data, used to obtain the  $\pi^0$  signal at  $p_T < 4$  GeV/c, all reconstructed BEMC points were used, even when they did not contain SMD clusters.

The SMD efficiency decreases rapidly and its energy resolution becomes poor with decreasing energy of the traversing particle, leading to significant fluctuations in the strip readout for  $E\lesssim 2$  GeV. Therefore, in the HighTower-1 data the SMD clusters were accepted only when they contained signals from at least two strips. This cut rejected a large fraction of the distorted and falsely split SMD clusters and reduced a possible effect of poor SMD response simulation at low energies.

# G. Charged particle veto using TPC

A charged particle veto (CPV) cut was applied to reject the charged hadrons that were detected in the calorimeter. A charged hadron was recognized as a BEMC cluster with a TPC track pointing to it. The cluster was rejected if the distance D between the BEMC point and the closest TPC track in the  $\eta$ - $\varphi$  coordinates was

$$D = \sqrt{(\Delta \eta)^2 + (\Delta \varphi)^2} < 0.04. \tag{2}$$

When a track was projected to the calorimeter surface, at a radius  $R=220\,\mathrm{cm}$ , this cut corresponded to a linear separation  $RD\approx 10\,\mathrm{cm}$  in the pseudorapidity range of this measurement. The efficiency of this cut was 35% in the MinBias data and 71% in the HighTower-1 and HighTower-2 data in all p+p and  $d+\mathrm{Au}$  datasets. The BEMC points remaining after this cut were considered to be photon candidates and were combined into pairs, defining the set of  $\pi^0$  candidates.

This veto introduced a false rejection of photon clusters if an unrelated charged particle happened to hit the calorimeter close

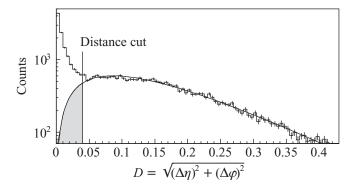


FIG. 4. Distribution of the distances D between BEMC points and their closest tracks, obtained from p+p HighTower-1 data in the bin  $4 < p_T < 5$  GeV/c. The curve shows a fit to Eq. (3); the vertical line indicates the CPV cut.

to the cluster. Figure 4 shows the distribution of D observed in the p+p data. In this plot, one distinguishes the peak of real charged particles at small distances, superimposed on a random component seen as a shoulder at larger distances. Assuming a uniform distribution of track projections in  $\eta$  and  $\varphi$  around the BEMC point, the radial distribution is given by

$$f(D) = De^{-D\rho}. (3)$$

Here  $\rho$  is the charged track density in the vicinity of the photon. This parameter was obtained from a simultaneous fit to the data in all bins of the event multiplicity M measured in the TPC, assuming a linear dependence on M,  $\rho = a + bM$ . The parametrization given by Eq. (3) describes well the random component, as shown by the curve in Fig. 4. The relative number of random coincidences that were falsely rejected was obtained by integrating the fitted curve up to the distance cut and weighting with the multiplicity distribution observed in each  $p_T$  bin. The resulting correction factor was  $\varepsilon_{\rm cpv}^\pi = 0.94 \pm 0.02$  for the p + p data and  $0.89 \pm 0.02$  for the p + p data.

In the direct photon analysis, the purity of the photon candidate sample was more important than in the  $\pi^0$  analysis; therefore, a stronger cut RD < 15 cm was used. The correction factors were calculated to be  $\varepsilon_{\rm cpv}^{\gamma} = 0.95 \pm 0.02$  for p+p and  $0.93 \pm 0.02$  for  $d+{\rm Au}$  data. The residual contamination by charged particles  $(C_{\pm})$  was estimated from the integrated excess of the D distribution over the fit to the random associations in the interval 15 < RD < 25 cm and was less than 5% for all  $p_T$  bins.

The uncertainties of these corrections contributed to a  $p_T$ -independent systematic uncertainty of the  $\pi^0$ ,  $\eta$ , and direct photon yields.

#### H. Photon conversions

A separate study was done to determine the degree to which the GEANT geometry described the distribution of material in the real STAR detector, and the corresponding correction factors  $c_{\rm loss}$  were extracted to account for any differences.

The photon conversion probability  $P_{\text{conv}}$  as a function of the depth d traversed in a material is given by

$$P_{\text{conv}} = 1 - \exp(-d/d_0),$$
 (4)

where  $d_0$  is the mean free path of the photon in that material. The probability that a  $\pi^0$  was not detected because at least one of its decay photons has converted is

$$P_{\text{loss}}^{\pi} = 2P_{\text{conv}}(1 - P_{\text{conv}}) + P_{\text{conv}}^{2}.$$
 (5)

The  $\pi^0$  losses due to conversions were in principle taken into account in the simulations mentioned in the sections that follow because the material traversed by the photons was included in the GEANT model of the detector. However, it was observed that the simulation failed to reproduce the number of photon conversions in the inner tracking system (SVT, SSD) and in the TPC inner field cage (IFC), all of which have a very complicated geometry of silicon sensors, readout electronics, and support structures [20,21,48]. The number of conversions in the simulated SVT, SSD, and IFC were underestimated by factors of  $\kappa = 2$ , 2, and 1.2, respectively, compared to that in the real data [49]. In simulations, the photon conversion probability in these detectors was in the range  $P_{\text{conv}} = 0.3\%$ –3.3%. To account for the missing material in the GEANT model, the photon spectra were corrected by factors  $c_{\rm loss}^{\gamma} = (1 - \sum P_{\rm conv})/(1 - \sum \kappa P_{\rm conv})$ , with the values of  $1.06 \pm 0.02$  and  $1.03 \pm 0.02$  for the p + p and  $d + {\rm Au}$ data, respectively. Using Eq. (5), this corresponds to correction factors of  $c_{\rm loss}^{\pi}=1.12\pm0.03~(p+p)$  and  $1.07\pm0.03~(d+{\rm Au})$  for the  $\pi^0$  spectra. Because the photon attenuation length in most absorbers rapidly approaches a constant for energies larger than  $\approx$ 100 MeV, the correction factors were assumed to be independent of the photon  $p_T$ .

# IV. NEUTRAL PION AND $\eta$ MESON ANALYSIS

The  $\pi^0$  and  $\eta$  were identified by their decays

$$\pi^0 \to \gamma \gamma$$
 and  $\eta \to \gamma \gamma$ .

These decay modes have branching ratios of 0.988 and 0.392, respectively [50]. The  $\pi^0$  lifetime is  $\tau = 8.4 \times 10^{-17}$  s, corresponding to a decay length  $c\tau = 0.025~\mu\text{m}$ . The  $\eta$  lifetime is even shorter (7 × 10<sup>-19</sup>s). Therefore, we assumed that the decay photons originated from the primary vertex. For each event, the invariant mass

$$m_{\gamma\gamma} = \sqrt{2E_1 E_2 (1 - \cos \psi)} \tag{6}$$

was calculated for all pairs of photons detected in the BEMC. Here  $E_1$  and  $E_2$  are the energies of the decay photons and  $\psi$  is the opening angle in the laboratory system. The reconstructed masses were accumulated in invariant mass spectra, where the  $\pi^0$  and the  $\eta$  showed up as peaks around their nominal masses  $(m_{\pi^0}=0.135~{\rm GeV}/c^2~{\rm and}~m_\eta=0.547~{\rm GeV}/c^2)$ . These peaks were superimposed on a broad distribution of combinatorial background, which originated from photon pairs that were not produced by the decay of a single parent particle.

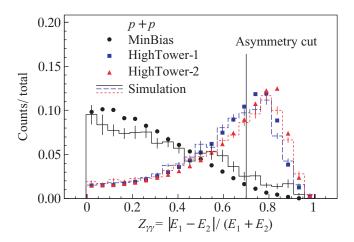


FIG. 5. (Color online) The energy asymmetry  $Z_{\gamma\gamma}$  of photon pairs reconstructed in p+p data (symbols) and in Monte Carlo simulation (histograms) for various triggers, normalized to unity for each trigger.

# A. Asymmetry of photon pairs

The energy asymmetry of the two-body decay of neutral mesons is defined as

$$Z_{\gamma\gamma} \equiv \frac{|E_1 - E_2|}{E_1 + E_2}. (7)$$

From the decay kinematics it follows that the probability for a given  $Z_{\nu\nu}$  is independent on  $Z_{\nu\nu}$ . Figure 5 shows the distribution of the asymmetry of photon pairs reconstructed in p + p data, including both  $\pi^0$  and  $\eta$  signals and background. In the MinBias data the distribution is not flat because of the acceptance effects—photons from an asymmetric decay have a large opening angle and one of them is likely to escape the BEMC. It is also seen that the HighTower energy threshold biases the asymmetry toward larger values because it is easier for an asymmetric decay to pass the trigger. The corresponding asymmetry distributions obtained from the Monte Carlo simulation, which represented the pure  $\pi^0$  signal, are also shown. The details of the simulation are given in Sec. IV G. Asymmetries observed in the simulation are in general agreement with those in the real data, considering the presence of background in the data.

In this analysis, the  $\pi^0$  and  $\eta$  candidates were only accepted if the asymmetry was less than 0.7. This cut rejected very asymmetric decays, where one of the BEMC points had low energy. It also rejected a significant part of the low-mass background (this background is described in Sec. IV D). The asymmetry cut improved the  $\pi^0$  signal-to-background ratio by a factor of  $\approx 1.5$ .

# **B.** Kinematic distributions

For each  $\pi^0$  candidate, the pseudorapidity  $\eta$ , the azimuth  $\varphi$ , the transverse momentum  $p_T$ , and the invariant mass  $m_{\gamma\gamma}$  [Eq. (6)] were calculated. Figure 6 shows the  $\eta$ ,  $\varphi$ ,  $p_T$ , and  $m_{\gamma\gamma}$  distributions of the  $\pi^0$  candidates in the p+p data. For the d+Au data these distributions are similar to those shown for p+p. The corresponding  $\eta$ ,  $\varphi$ , and  $p_T$  distributions of the  $\pi^0$ 's reconstructed in the simulation are also shown.

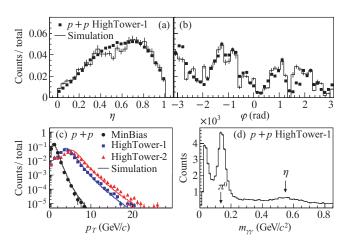


FIG. 6. (Color online) Distribution of  $\pi^0$  candidates obtained from the p+p data, as a function of (a)  $\eta$  and (b)  $\varphi$ , and (c)  $p_T$  and (d)  $m_{\gamma\gamma}$ .

The  $\eta$  distribution shows the decrease of the calorimeter acceptance at  $\eta=0$  and  $\eta=1$ , because it is likely that one of the decay photons at the calorimeter edges escapes detection. The asymmetry of the  $\eta$  distribution is attributable to the fact that the calorimeter half barrel is positioned asymmetrically with respect to the interaction region. The azimuthal dependence of the calorimeter acceptance was caused by failing SMD modules (the data used in this article are from the early years of detector operation, in which such failures occurred relatively frequently), as well as by dead and hot towers. The gross features of the data reflecting the calorimeter acceptance are reasonably well reproduced by the pure  $\pi^0$  Monte Carlo simulations.

Figure 6(c) shows the  $p_T$  distribution of the photon pairs, separately for the MinBias and for the two HighTower datasets. The HighTower trigger threshold effects are reasonably well reproduced in the simulation. It is seen that using the HighTower triggers significantly increased the rate of  $\pi^0$  candidates at high  $p_T$ .

The  $p_T$ -integrated invariant mass distribution in Fig. 6(d) clearly shows the  $\pi^0$  and  $\eta$  peaks superimposed on a broad background distribution. This background has a combinatorial and a low-mass component, discussed in detail in the two following sections.

# C. Combinatorial background

The combinatorial background in the invariant mass distribution originated from pairs of photon clusters that were not produced by a decay of a single particle. To describe the shape of the combinatorial background, we used the event-mixing technique, where photon candidates from two different events were combined. To avoid the mixing of different event topologies, the data were subdivided into mixing classes based on the vertex position, BEMC multiplicity, and trigger type (MinBias, HighTower-1, and HighTower-2).

Figure 7(a) shows, as an example, the invariant mass distribution in the  $4 < p_T < 5$  GeV/c bin, obtained from the p + p HighTower-1 data, together with the combinatorial

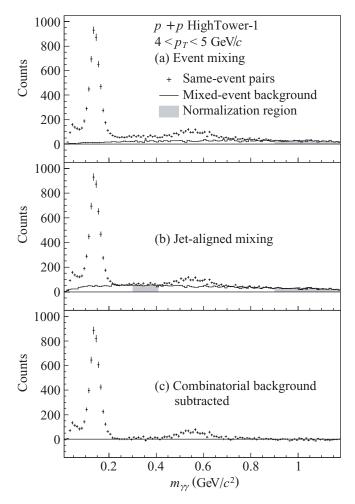


FIG. 7. The same-event invariant mass distribution (crosses) and combinatorial background (histogram) observed in one  $p_T$  bin of the p+p data. Background estimated (a) from random event mixing and (b) from a linear combination of random and jet-aligned mixing; (c) background-subtracted distribution. The shaded areas indicate the regions where the mixed-event background was normalized to the same-event distributions.

background obtained from the event mixing. The mixedevent background distribution was normalized to the sameevent distribution in the invariant mass region  $0.9 < m_{\gamma\gamma} < 1.2 \text{ GeV}/c^2$ .

There is still some residual background in the interval  $0.2 < m_{\gamma\gamma} < 0.4 \text{ GeV}/c^2$ . This background is attributable to correlation structures (jet structures) in the event, which are not present in the sample of mixed events. To preserve jet-induced correlations, the jet axes in both events were aligned before mixing [38], as described later in this article.

To determine the  $\eta$ - $\varphi$  position of the most energetic jet in every event, the cone algorithm was used [51]. The mixed-event  $\pi^0$  candidates were constructed by taking two photons from different events, where one of the events was displaced in  $\eta$  and  $\varphi$  by  $\Delta \eta = \eta_2 - \eta_1$  and  $\Delta \varphi = \varphi_2 - \varphi_1$ , respectively. Here  $\eta_{1,2}$  and  $\varphi_{1,2}$  are the jet orientations in the two events.

Figure 8 shows a schematic view of two superimposed events where the jet axes are aligned. To minimize acceptance distortions, the events were divided into ten mixing classes in

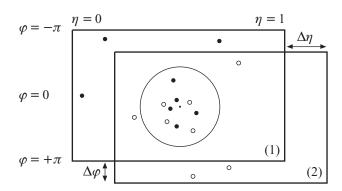


FIG. 8. A schematic view of two superimposed events, where the jet axes are aligned to preserve the jet-induced correlations in the mixed event.

the jet  $\eta$  coordinate. By mixing only events in the same class, the shift  $\Delta \eta$  was limited to 0.1. Because the calorimeter has a cylindrical shape, the shift in  $\varphi$  did not induce any significant acceptance distortion.

A side effect of this procedure was that correlations were induced if there was no real jet structure, because the jet-finding algorithm then simply picked the most energetic track in the event. To reduce possible bias introduced by such correlations, the combinatorial background was taken to be fully random for  $p_T < 1.2 \text{ GeV}/c$  and fully jet-aligned for  $p_T > 10 \text{ GeV}/c$ . Between these values, the random component decreased linearly with increasing  $p_T$ . We assigned a systematic uncertainty of 10% to the random background fraction, which resulted in a systematic uncertainty of 5% of the  $\pi^0$  and 3.5% of the  $\eta$  yields.

Figure 7(b) shows the same invariant mass spectrum as that shown in Fig. 7(a), with the background estimated by the combined random and jet-aligned event mixing. The mixed-event background was normalized to the same-event distribution in the ranges  $0.3 < m_{\gamma\gamma} < 0.4$  and  $0.9 < m_{\gamma\gamma} < 1.2~{\rm GeV}/c^2$ . By changing the subtracted background within its normalization uncertainty, we obtained another component of a systematic error of the  $\pi^0$  and  $\eta$  yields, which was found to increase with increasing  $p_T$  from 0.5% to 3% for the  $\pi^0$  and from 10% to 50% for the  $\eta$ .

Figure 7(c) shows the background-subtracted distribution. It is seen that there is still a residual background component at invariant mass  $m_{\gamma\gamma} < 0.1~{\rm GeV}/c^2$ . The origin of this background is described in the next section.

# D. Low-mass background

Random fluctuations in the SMD signals occasionally generate a double-peaked hit structure, in which case the clustering algorithm incorrectly splits the cluster. These random fluctuations enhance the yield of pairs with minimal angular separation and thus contribute to the lowest diphoton invariant mass region, as can be seen in Fig. 7(c). However, at a given small opening angle, the invariant mass increases with increasing energy of the parent particle, so that the low-mass background distribution extends to larger values of  $m_{\gamma\gamma}$  with increasing photon  $p_T$ .

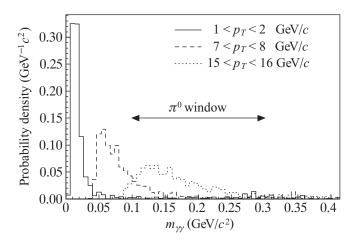


FIG. 9. The simulated low-mass background distributions from erroneous splitting of single photons in three bins of the reconstructed pair  $p_T$ . The distributions extend to larger invariant masses with increasing  $p_T$  and move into the  $\pi^0$  region (shown for  $p_T = 15 \text{ GeV}/c$ ).

The shape of the low-mass background was obtained from a simulation as follows. Single photons were generated with flat distributions in  $-\pi < \varphi < +\pi$ ,  $-0.2 < \eta < 1.2$ , and  $0 < p_T < 25$  GeV/c. These photons were tracked through a detailed description of the STAR geometry with the GEANT program. A detailed simulation of the electromagnetic shower development in the calorimeter was used to generate realistic signals in the towers and in the SMD. The simulated signals were processed by the same reconstruction chain as the real data. Photons with more than one reconstructed cluster were observed, and  $m_{\gamma\gamma}$  and  $p_T$  of such cluster pairs were calculated. The  $m_{\gamma\gamma}$  histograms were accumulated, with each entry weighted by the  $p_T$  spectrum of photons in the real data, corrected for the photon detection efficiency.

Figure 9 shows the low-mass background distributions in three bins of the reconstructed pair  $p_T$ . It is seen that the distributions indeed move to larger invariant masses with increasing  $p_T$  and extend far into the  $\pi^0$  window at high  $p_T$ . For this reason, it was not possible to estimate the amount of this background from a phenomenological fit to the data, and we had to rely on the Monte Carlo simulation to subtract this background component.

The second significant source of BEMC clusters that passed the CPV cut was the neutral hadrons produced in the collisions, mostly antineutrons above 2 GeV/c. To account for the additional low-mass background from these hadrons, simulations of antineutrons were performed in the same way as those of photons, and the reconstructed invariant mass distribution was added according to the realistic proportion  $\bar{n}/\gamma_{\rm incl}$ , where the antineutron yield was estimated as described in Sec. VB.

The low-mass background was normalized by matching the observed  $p_T$  spectrum of the clusters between simulation and data. This removal procedure worked well, and Fig. 10 shows the invariant mass spectra and the low-mass background component, and the final background-subtracted spectrum for the p + p HighTower-1 data. The normalization uncertainty

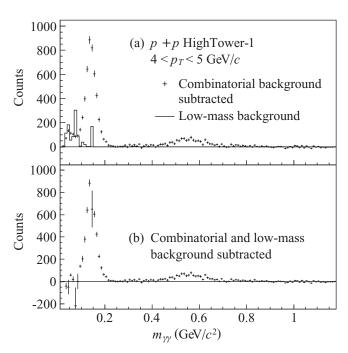


FIG. 10. The invariant mass distribution observed in one  $p_T$  bin of the p + p data (a) before and (b) after the low-mass background subtraction.

of the low-mass background contributes to the systematic uncertainty of the  $\pi^0$  cross section and reaches 15% at the high- $p_T$  end of the spectrum.

# E. Peak position and width

Figure 11(a) shows the background-subtracted  $m_{\nu\nu}$  distribution in the region  $4 < p_T < 5 \text{ GeV}/c$  obtained from the p + p HighTower-1 data (symbols), together with the corresponding distribution from the detector simulation (histogram). To compare the real and simulated  $m_{\nu\nu}$  distributions for all bins in  $p_T$  and for all datasets, we estimated the position and width of the peaks using Gaussian fits in the peak region. Figure 11(b) shows the peak positions obtained from the fit to the p + p data. It is seen that the peak position shifts toward higher masses with increasing  $p_T$ . This shift is a manifestation of the bin migration effect that originates from statistical fluctuations in the calorimeter response. Due to the steeply falling  $p_T$  spectrum, the energy resolution causes a net migration toward larger  $p_T$ . Because larger values of  $p_T$  imply larger values of  $m_{\gamma\gamma}$ , the migration effect biases the invariant mass peak toward larger values.

An additional peak shift at the largest values of  $p_T$  is caused by the SMD strip granularity, which imposes a lower limit on the opening angle of the reconstructed photon pairs. The minimal SMD cluster separation in each dimension that can be resolved by the cluster finder is 1.5 strips, and most clusters contain at least two or three strips. Therefore, the pair reconstruction is less efficient for the symmetric decays with the smallest opening angles at  $p_T \gtrsim 10~{\rm GeV}/c$ . This leads to an increased average opening angle and  $m_{\gamma\gamma}$  of the reconstructed photon pairs from  $\pi^0$  decays.

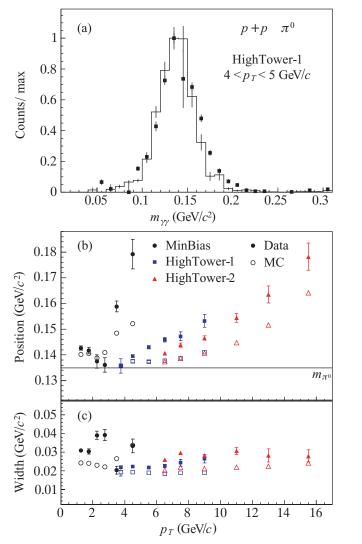


FIG. 11. (Color online) (a) Invariant mass spectrum of  $\pi^0$ 's reconstructed in the simulation (histogram) in comparison to the p+p HighTower-1 data (symbols) in the  $4 < p_T < 5$  GeV/c bin. (b) Peak position and (c) width in the real data (solid symbols) and in the simulation (open symbols). The horizontal line in panel (b) shows the true  $\pi^0$  mass.

The peak position observed in the data is larger than that found in the simulations by  $(3.5\pm0.6)\%$ , on average. This difference could be caused by the global energy scale of the BEMC towers being off by a similar amount. We already accounted for this possibility by assigning a systematic uncertainty of 5% to the BEMC calibration constants (see Sec. III E).

Figure 11(c) compares the  $\pi^0$  peak width in the data and in the simulation, and it is seen that the peak width in the data is larger than that in the simulation by  $(25 \pm 2)\%$ , on average. This is a sufficiently good agreement for this analysis, because the  $\pi^0$  and  $\eta$  yields were counted in the mass windows that were adjusted in each  $p_T$  bin to cover the entire signal peak.

The peak shape of the  $\eta$  meson, as well as its position and width, are shown in Fig. 12, as a function of the reconstructed  $p_T$ . The peak position in the data is larger than in the

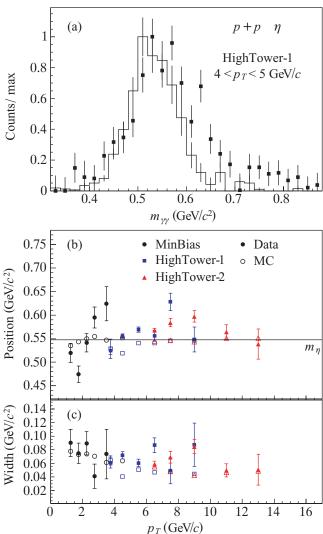


FIG. 12. (Color online) (a) Invariant mass spectrum of  $\eta$  mesons reconstructed in the simulation (histogram) in comparison to the p+p HighTower-1 data (symbols) in the  $4 < p_T < 5$  GeV/c bin. (b) Peak position and (c) width in the real data (solid symbols) and in the simulation (open symbols). The horizontal line in panel (b) shows the true  $\eta$  mass.

simulations by  $(5.1\pm1.2)\%$ , and the width by  $(34\pm11)\%$ , on average. The difference in the peak position is similar to the preceding  $\pi^0$  case, which supports the possibility of both being caused by a small systematic offset in the BEMC calibration. The observed level of agreement between data and simulations is considered to be sufficient for this analysis.

# F. Invariant yield extraction

The invariant yield of the  $\pi^0$  and  $\eta$  mesons per MinBias collision, as a function of  $p_T$ , is given by

$$E\frac{d^3N}{d\mathbf{p^3}} = \frac{d^3N}{p_T dp_T dy d\varphi} = \frac{d^2N}{2\pi p_T dp_T dy},\tag{8}$$

where in the last equality integration over the full  $2\pi$  azimuthal coverage of the STAR detector is performed. Using the

experimentally measured quantities, the invariant yield was calculated as

$$E\frac{d^{3}N}{d\mathbf{p}^{3}} = \frac{1}{2\pi p_{T} N_{\text{trig}} K_{\text{trig}}} \frac{Y}{\Delta p_{T} \Delta y} \frac{\varepsilon_{\text{vert}} c_{\text{loss}}^{\pi}}{\varepsilon_{\text{acc}}^{\pi} \varepsilon_{\text{cDV}}^{\pi}} \frac{1}{\mathcal{B}_{\gamma\gamma}}, \quad (9)$$

where

- (i) Y is the raw yield measured in the bin centered at p<sub>T</sub> and y;
- (ii)  $\Delta p_T$  is the width of the  $p_T$  bin for which the yield was calculated;
- (iii)  $\Delta y$  is the rapidity range of the measurement; in this analysis  $\Delta y = 1$  for all data points, except for the  $\eta$  yields at  $p_T < 3 \text{ GeV}/c$ , where the correction for the difference between rapidity and pseudorapidity reached 7%.
- (iv)  $N_{\text{trig}}$  is the number of triggers recorded;
- (v)  $K_{\text{trig}}$  is the trigger scale factor;  $K_{\text{trig}} \equiv 1$  for the MinBias events and >1 for the HighTower events; the product  $N_{\text{trig}}K_{\text{trig}}$  is the equivalent number of MinBias events that produced the yield Y;
- (vi)  $\varepsilon_{\text{vert}}$  is the vertex finding efficiency in MinBias events;
- (vii)  $c_{\text{loss}}^{\pi}$  is the correction for the missing material in the simulation;
- (viii)  $\varepsilon_{\rm acc}^{\pi}$  is the BEMC acceptance and efficiency correction factor:
- (ix)  $\varepsilon_{\text{cpv}}^{\pi}$  is a correction for random TPC vetoes;
- (x)  $\mathcal{B}_{\gamma\gamma} = \Gamma_{\gamma\gamma}/\Gamma$  is the branching ratio of the diphoton decay channel (0.988 for  $\pi^0$  and 0.392 for  $\eta$  [50]).

The raw  $\pi^0$  and  $\eta$  yields were counted in the  $p_T$ -dependent  $m_{\gamma\gamma}$  windows that contained the peaks. The low-mass border of the  $\pi^0$  peak region was taken to be a linear function of  $p_T$ , common for all datasets and triggers. This cut was optimized to capture most of the yield and as little low-mass background as possible. The high-mass border also linearly increased with  $p_T$  to cover the asymmetric right tail of the peak. Similarly, the  $\eta$  peak region was a  $p_T$ -dependent window that captured most of the signal. For completeness, we give the following parametrization of the  $\pi^0$  and  $\eta$  windows:

$$75 + 1.7p_T < m_{\gamma\gamma}(\pi^0) < 250 + 3.3p_T \text{ MeV}/c^2,$$
  
 $350 + 3.3p_T < m_{\gamma\gamma}(\eta) < 750 \text{ MeV}/c^2,$  (10)

where  $p_T$  is measured in GeV/c. The stability of the yields was determined by varying the vertex position cut, the energy asymmetry cut, and the yield integration windows. From the observed variations, a point-to-point systematic error of 5% was assigned to the  $\pi^0$  and  $\eta$  yields.

Within each trigger in the p+p data, the  $\pi^0$  signal significance decreased from  $\approx 34$  to  $\approx 6$  standard deviations with increasing  $p_T$ , because of the corresponding reduction in statistics. In the d+Au data, the same trends were observed, but the significance was lower than in the p+p data by a factor of 1.9, on average, which is mainly caused by the lower integrated nucleon-nucleon luminosity in these data. The significance of the  $\eta$  signal was between 18 and 2.5 standard deviations in the p+p data and between 5.5 and 1.0 standard deviations in the d+Au data.

# G. Acceptance and efficiency correction

To calculate the detector acceptance and reconstruction efficiency correction factor  $\varepsilon_{\rm acc}^\pi$ , a Monte Carlo simulation of the detector was used. The  $\pi^0$  decay photons were tracked through the STAR detector geometry using GEANT. The simulated signals were passed through the same analysis chain as the real data.

The  $\pi^0$ 's were generated in the pseudorapidity region  $-0.3 < \eta < +1.3$ , which is sufficiently large to account for edge effects caused by the calorimeter acceptance limits of  $0 < \eta < 1$ . The azimuth was generated flat in  $-\pi < \varphi < +\pi$ . The  $p_T$  distribution was taken to be uniform up to 25 GeV/c, which amply covers the measured pion  $p_T$  range of up to 17 GeV/c. The vertex distribution of the generated pions was taken to be Gaussian in z with a spread of  $\sigma = 60$  cm and centered at z = 0.

The generated  $\pi^0$ 's were allowed to decay into two photons,  $\pi^0 \to \gamma \gamma$ . The GEANT simulation accounted for all interactions of the decay photons with the detector, such as conversion into  $e^+e^-$  and showering in the calorimeter or in the material in front of it.

To reproduce a realistic energy resolution of the calorimeter, an additional smearing had to be applied to the energy deposits calculated by GEANT. In all simulations, a spread of 5% was used to reproduce the p+p data and 10% for the d+Au data.

To reproduce the  $p_T$  spectrum of pions in the data, each Monte Carlo event was weighted by a  $p_T$ -dependent function. This weighting technique allowed us to sample the entire  $p_T$  range with good statistical power, while reproducing the bin migration effect caused by the finite detector energy resolution. An NLO pQCD calculation [52] provided the initial weight function, which was subsequently adjusted in an iterative procedure.

The time dependence of the calorimeter acceptance during data taking was recorded in database tables that were used in the analysis. To reproduce this time dependence in the simulation, the generated events were assigned time stamps that followed the timeline of the real data taking. In this way, the geometrical acceptance of the calorimeter (mean fraction of good towers) was reproduced in the Monte Carlo with a precision of  $\approx 0.5\%$ .

In the analysis of real data, we used vertices reconstructed from TPC tracks and those derived from BBC time-of-flight measurements. The former have submillimeter resolution, whereas the latter have a precision of only  $\approx$ 40 cm. To account for the BBC vertex resolution, 35% of the generated pions in the p+p MinBias data had their point of origin artificially smeared in the z direction. No such smearing was applied to the other simulated data, where BBC vertex was not used.

The acceptance and efficiency correction factor was calculated from the simulation as the ratio of the raw  $\pi^0$  yield reconstructed in a  $p_T$  bin to the number of simulated pions with the true  $p_T$  generated in that bin. This was done separately for each trigger, using the same  $\pi^0$  reconstruction cuts as was done in the real data analysis. In particular, the reconstructed value of pseudorapidity was required to fall in the range  $0 < \eta < 1$  in both the data and the simulation, while in the latter the generated value of  $\eta$  was also required to fall in this range.

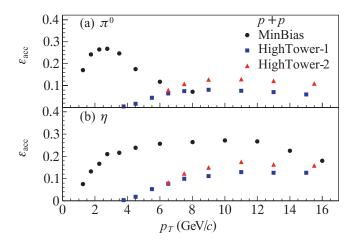


FIG. 13. (Color online) Acceptance and efficiency correction factor  $\varepsilon_{\rm acc}$  for (a)  $\pi^0$  and (b)  $\eta$  calculated from the Monte Carlo simulation for the p+p data.

As an example, Fig. 13 shows the  $\pi^0$  and  $\eta$  correction factors for the three triggers in p + p data.

The difference between the MinBias and HighTower correction factors was caused by the SMD requirement in the HighTower data, which was absent in the MinBias data. The absence of the SMD information reduced the  $\pi^0$  reconstruction efficiency in the MinBias data at  $p_T > 3 \text{ GeV}/c$ , where the decay photons were separated by less than two towers. The  $\eta$  reconstruction is only affected by this at larger values of  $p_T$ .

The effect of the SMD quality requirement of having at least two adjacent strips in a cluster is illustrated in Fig. 14, which shows the correction factor calculated for the p+p HighTower-1 data with (solid symbols) and without (open symbols) the SMD quality requirement. This requirement reduced the number of accepted  $\pi^0$  candidates by  $\approx$ 45%. This explains the difference between the HighTower-1 and HighTower-2 (no SMD quality cut) correction factors at high  $p_T$  seen in Fig. 13.

The current simulation framework poorly reproduces the shower shapes in the SMD at the low incident photon energies. To account for residual bias after applying the SMD quality cut, we assigned a systematic uncertainty to the HighTower-1 cross section, which decreases from 15% at  $p_T = 4 \text{ GeV}/c$  to zero at  $p_T = 7 \text{ GeV}/c$ .

To determine a dependence of the acceptance correction on the track multiplicity M, and thus on the centrality, we analyzed a sample of generated  $\pi^0$ 's embedded into real

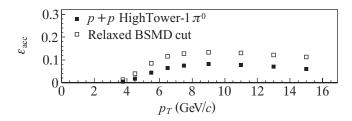


FIG. 14. Acceptance and efficiency correction factor  $\varepsilon_{\rm acc}$  for the p+p HighTower-1 data, with the standard set of cuts (solid symbols) and with the SMD quality cut removed (open symbols).

 $d+{\rm Au}$  events. No significant centrality dependence was found. Therefore, the same correction factors were applied to the different centrality classes in the  $d+{\rm Au}$  data. The dependence of the efficiency on the locally higher multiplicity in jets was investigated in a PYTHIA [53] simulation, and no significant difference in the efficiency was observed relative to a single-particle simulation.

# H. HighTower trigger normalization

We have shown in Fig. 6 the  $p_T$  distribution of  $\pi^0$  candidates for the p+p MinBias, HighTower-1, and HighTower-2 data. To normalize the HighTower spectra to those of the MinBias,  $p_T$ -independent scale factors were applied. These scale factors were estimated as the ratio of observed MinBias to HighTower event rates,

$$K_{\text{trig}} = \sum N_{\text{MB}} S_{\text{MB}} / \sum N_{\text{HT}} S_{\text{HT}}.$$
 (11)

Here  $N_{\rm MB}$  and  $N_{\rm HT}$  are the numbers of MinBias and High-Tower triggers that passed the selection cuts,  $S_{\rm MB}$  and  $S_{\rm HT}$  are the online prescale factors adjusted on a run-by-run basis to accomodate the DAQ bandwidth, and the sums are taken over all runs where both the MinBias and the HighTower triggers were active. We obtained the values  $K_{\rm trig} = 4.67 \times 10^3$  and  $1.96 \times 10^4$  for the p+p HighTower-1 and HighTower-2 triggers, respectively, and  $K_{\rm trig} = 2.87 \times 10^3$  and  $2.86 \times 10^4$  for the d+Au triggers.

To check the scale factors, the HighTower software filter, which simulated the hardware trigger, was applied to the MinBias data. The scale factors were obtained as the ratio of the total number of MinBias events to the number that passed the filter. To obtain a more precise HighTower-1/HighTower-2 relative normalization factor, the software filter was applied to the HighTower-1 data. The results from the two methods agreed within 3% for HighTower-1 data and within 5% for HighTower-2 data. These numbers were taken as the systematic uncertainties of the HighTower normalization factors.

The difference between vertex finding efficiencies in Min-Bias and HighTower data was effectively absorbed in the scale factor  $K_{\rm trig}$ . Therefore, the vertex-finding efficiency correction was applied to the scaled HighTower data, as well as to the MinBias data.

# I. Bin-centering scale factors

To assign the yield measured in a  $p_T$  bin to a single  $p_T$  value, the procedure from Ref. [54] was applied. The variation of the yield within a bin was approximated by the function  $f(p_T) = A \exp(-Bp_T)$ . The measured yield in the bin was assigned to the momentum  $p_T^*$  calculated from the equation

$$f(p_T^*) = \frac{1}{\Delta p_T} \int_{\Delta p_T} f(p_T) dp_T.$$
 (12)

The procedure was repeated, taking  $p_T^*$  as the abscissa, until the  $p_T^*$  values were stable (typically after three iterations).

To facilitate the comparison of results from the various datasets, the yields were scaled to the bin centers by the ratio  $K_{\text{bin}} = f(p_T^*)/f(p_T)$ , where  $p_T$  is the center of the bin. The

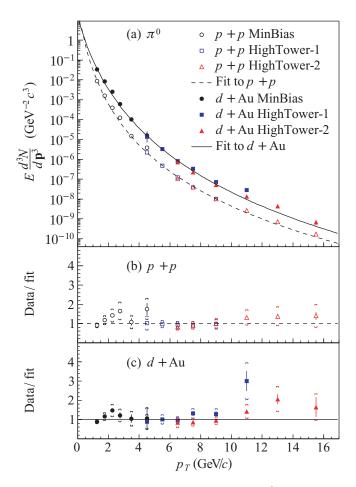


FIG. 15. (Color online) (a) Invariant yield of  $\pi^0$  per MinBias p+p and d+Au collision. Curves are the power law fits given in the text. Invariant yield divided by the fit to the (b) p+p and (c) d+Au data. The error bars are statistical and brackets in the lower panels are the systematic uncertainties.

statistical and systematic errors were also scaled by the same factor.

To estimate a systematic uncertainty introduced by this procedure, we changed the functional form of  $f(p_T)$  either to a local power law in each bin or to a global power law in the full  $p_T$  range. The observed variation in  $K_{\rm bin}$  was below 1.5% in most  $p_T$  bins.

# J. Fully corrected yields

The fully corrected  $\pi^0$  and  $\eta$  invariant yields per MinBias event [Eq. (9)] are shown in the top panels of Figs. 15 and 16, respectively. The curves represent power law fits to the data of the form

$$E\frac{d^3N}{d\mathbf{p}^3} = \frac{A}{(1+p_T/p_0)^n}. (13)$$

The ratios between the data and the fits are shown in panels (b) and (c) in both figures. The brackets in the lower panels are the systematic uncertainties (see Sec. IV K), which are partially correlated between different trigger datasets. The agreement between the data taken with the different

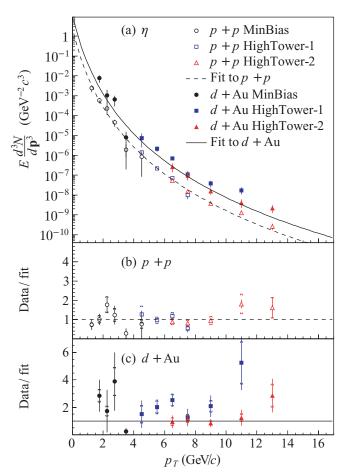


FIG. 16. (Color online) (a) Invariant yield of  $\eta$  per MinBias p+p and d+Au collision. Curves are the power law fits given in the text. Invariant yield divided by the fit to the (b) p+p and (c) d+Au data. The error bars are statistical and brackets in the lower panels are the systematic uncertainties.

triggers is satisfactory, although a small systematic difference between the d + Au HighTower-1 and HighTower-2 yields (1.9 standard deviations, on average) was observed.

For the calculation of the final cross sections and their ratios, the data from three triggers were merged. The HighTower-1 data were used in the MinBias-HighTower-1 overlap bins, because the MinBias data in those bins almost entirely represent a subset of HighTower-1 events, selected by an online prescale factor  $S_{\rm MB}$ . Similarly, the HighTower-2 data were used in the HighTower-1-HighTower-2 overlap bins.

## K. Summary of systematic uncertainties

The uncertainty of the calorimeter tower calibration was the dominant source of systematic uncertainty in this analysis. The uncertainty of the uncorrected yield  $Y(p_T)$  was estimated from

$$\delta Y(p_T) = \left| \frac{dY(p_T)}{dp_T} \right| \delta p_T, \tag{14}$$

where  $\delta p_T$  was taken to be 5% in the d + Au and p + p data (as derived from the electron calibration; see Sec. III E). This

TABLE III. Systematic error contributions. The classifications A, B, C, and N are defined in the text. The error
contributions to the $\pi^0$ cross section, the $\eta/\pi^0$ ratio, $R_{CP}$ , $R_{dA}$ , and the inclusive photon yield are indicated in the
respective columns.

Source	Type	Value at low (high) $p_T$ (%)	$Ed^3\sigma/d\mathbf{p^3}$	$\eta/\pi^0$	$R_{CP}$	$R_{dA}$	$\gamma_{ m incl}$
Combinatorial background	A	0.5 (3)	+	+	+	+	
Mixed-event background	C	5	+	+			
Low-mass background	C	1 (15)	+	+			
Random vetoes	N	2	+		+	+	+
HighTower normalization	В	3 (5)	+		+	+	+
Analysis cuts	A	5	+	+	+	+	
Conversion correction	В	3 for $\pi^0$ and $\eta$ , 2 for $\gamma_{incl}$	+			+	+
Tower energy scale	В	15 (35) in $p + p$ , 10 (40) in $d + Au$	+			+	+
SMD simulation	C	15 (0) at $p_T = 4(7) \text{ GeV}/c$	+				
SMD energy scale	В	3.5 (1) at $p_T = 4(11) \text{ GeV}/c$	+			+	
Bin centering	C	1.5	+				
Vertex finding efficiency	N	1 in $d$ + Au MinBias	+		+	+	+
MinBias cross section	N	11.5 in $p + p$ , 5.2 in $d + Au$	+			+	+
Glauber model $\langle N_{\rm coll} \rangle$	N	5.3 in $d$ + Au MinBias, 10.5 in $R_{CP}$			+	+	

 $p_T$ -dependent systematic uncertainty was, on average, 25% in both p+p and d+Au data.

Another strongly  $p_T$ -dependent term in Eq. (9) is the acceptance and efficiency correction factor  $\varepsilon_{\rm acc}$ , obtained from the Monte Carlo simulation. However, this term contributed much less to the cross-section uncertainty, because the simulations used the inverse of real calibration constants to convert the GEANT energy deposit in each tower to the ADC value. Therefore, the energies later reconstructed from the ADC values with the same constants were not sensitive to their fluctuations. This factor was sensitive to the tower calibration only through the HighTower threshold values and contributed a systematic uncertainty of 5% in the p+p data and 8% in the d+Au data, on average, correlated with the uncorrected yield uncertainty estimated previously.

The uncertainty of the SMD energy scale entered the analysis mainly due to threshold effects in the clustering, where the loss of a soft photon from the asymmetrically decayed  $\pi^0$  may change the reconstruction efficiency. This analysis does not depend on the absolute energy calibration of the SMD because the main energy measurement was obtained from the towers. Instead, we estimated the possible disagreement between the SMD scale in the data and in Monte Carlo simulation to be below 20%. The resulting variation of  $\varepsilon_{\rm acc}$  from this source was 3.5% in the HighTower data at the pion  $p_T=4~{\rm GeV}/c$  and less than 1% above 11  ${\rm GeV}/c$ .

All systematic error contributions are summarized in Table III and classified into the following categories:

A, point-by-point systematic uncertainty;

B,  $p_T$ -correlated systematic uncertainty, but uncorrelated between p + p and d + Au datasets;

C,  $p_T$ -correlated systematic uncertainty, also correlated between p + p and d + Au datasets;

N, normalization uncertainty, uncorrelated between p+p and  $d+\mathrm{Au}$  datasets.

Table III also lists which measurement is affected by a given source of systematic error.

#### V. DIRECT PHOTON ANALYSIS

The traditional approach to measuring direct photon production in hadronic collisions uses isolation criteria. Photons from decays of highly energetic hadrons should be accompanied by other jet fragments. Therefore, one can reject those by requiring less than a certain amount of background energy in a cone around a photon candidate [55]. However, prompt photon production beyond leading order in pQCD cannot be separated unambiguously from photons from fragmentation processes, although the framework for applying isolation cuts in pQCD calculations is established and theoretical interpretation of experimental results is possible. In addition, the use of isolation cuts in the high-multiplicity environment of heavy-ion collisions is not straightforward.

Because the present analysis is intended to provide a baseline measurement for heavy-ion collisions, we have chosen to use the method of statistical subtraction to obtain direct photon yields. For this method, one obtains inclusive photon spectra, which, in addition to the direct contribution, contain a large background of decay photons, dominantly from  $\pi^0$  decays. An accurate measurement of  $\pi^0$  and heavier hadrons provides the necessary input to subtract the decay background. This method has been successfully used in heavy-ion reactions [56,57]; however, it does not provide event-by-event direct photon identification.

The sample of photon candidates served as the main input to the direct photon analysis, as in the case of the reconstruction of the  $\pi^0$  spectrum, described in the previous sections. After subtracting the contamination by charged particles and neutral hadrons, the raw inclusive photon sample was corrected to account for the limited acceptance and the finite detector resolution. In parallel, the total yield of photons from  $\pi^0$ ,  $\eta$ , and  $\omega(782)$  decays was simulated, assuming a phenomenological scaling law ( $m_T$  scaling) for the  $\eta$  and  $\omega(782)$  spectra.

To exploit the fact that the inclusive photon and decay photon yields have many correlated uncertainties, we studied the direct photon yield via the double ratio

$$R_{\gamma} \equiv \frac{\gamma_{\rm incl}/\pi^0}{\gamma_{\rm decay}/\pi^0},\tag{15}$$

where the numerator equals the point-to-point ratio of the measured spectra of inclusive photons and  $\pi^0$ 's, as a function of  $p_T$ , and the denominator is the simulated background contribution from decay processes divided by the parametrized  $\pi^0$  yield. It follows that

$$R_{\gamma} = 1 + \frac{\gamma_{\text{dir}}}{\gamma_{\text{decay}}},\tag{16}$$

which serves as an indicator of a direct photon signal  $\gamma_{\rm dir}$  ( $R_{\gamma} > 1$ ). The absolute direct photon yields can subsequently be determined as

$$\gamma_{\rm dir} = \left(1 - R_{\gamma}^{-1}\right) \gamma_{\rm incl},\tag{17}$$

where the systematic uncertainties, which canceled in the double ratio, have to be included again.

# A. Inclusive photons

The reconstruction of the inclusive photon spectrum was in many ways similar to that of the  $\pi^0$ . The uncorrected photon spectrum was extracted from the same data sample using identical event and photon candidate cuts. However, in the  $\pi^0$  analysis there were no rigorous constraints on the purity of the photon candidates, because remaining contributions from charged particles and neutral hadrons were identified afterward as the combinatorial background in the mass distributions. In contrast, the uncorrected inclusive photon yield  $Y_{\rm incl}$  was obtained from an explicit subtraction of such backgrounds,

$$Y_{\text{incl}} = (1 - C_0)(1 - C_+)Y_{\text{cand}},$$
 (18)

where the correction terms  $C_0$  and  $C_{\pm}$  represent the fractional contamination by neutral hadrons and charged particles, respectively. The charged-particle contamination  $C_{\pm}$  was estimated in Sec. III G and found to be smaller than 5%.

The invariant yield of inclusive photons was calculated, similarly to that of  $\pi^0$ 's in Eq. (9), as

$$E\frac{d^{3}N}{d\mathbf{p}^{3}} = \frac{1}{2\pi p_{T} N_{\text{trig}} K_{\text{trig}}} \frac{Y_{\text{incl}}}{\Delta p_{T} \Delta y} \frac{\varepsilon_{\text{vert}} c_{\text{loss}}^{\gamma}}{\varepsilon_{\text{acc}}^{\gamma} \varepsilon_{\text{cpv}}^{\gamma}}.$$
 (19)

Here  $\varepsilon_{\rm acc}^{\gamma}$  is the single photon acceptance and efficiency correction factor, discussed in Sec. V C.

# B. Neutral hadron background

The term  $C_0$  in Eq. (18) was defined as the number of reconstructed showers generated by neutral hadrons relative to the total number of showers in the photon candidate sample. The STAR detector has no means of directly identifying neutrons and antineutrons. Therefore, this contamination was simulated using the measured (anti)proton spectra as input.

The largest source of neutral contamination was the  $\bar{n}$  annihilation in the calorimeter, for example  $\bar{n}+p\to 2\pi^+\pi^-\pi^0$ . This initiates a shower that does not necessarily develop in the incident direction of  $\bar{n}$ . Moreover, the available energy for the

reaction products includes twice the rest mass of a nucleon ( $\approx$ 2 GeV).

STAR has measured the p and  $\bar{p}$  production in p+p and  $d+{\rm Au}$  collisions [24]. The reported yields, however, were not corrected for the  $\Lambda$  and  $\bar{\Lambda}$  feed-down, which is expected to have a contribution of  $\delta_{\Lambda}\approx 20\%$  [23]. Therefore, the  $\bar{n}$  yield was estimated as

$$Y(\bar{n}) = (1 - \delta_{\Lambda}) Y(\bar{p}) + \delta_{\Lambda} \frac{\mathcal{B}(\Lambda \to n\pi^{0})}{\mathcal{B}(\Lambda \to p\pi^{-})} Y(\bar{p}), \quad (20)$$

where the branching ratios are  $\mathcal{B}(\Lambda \to n\pi^0) = 0.358$  and  $\mathcal{B}(\Lambda \to p\pi^-) = 0.639$  [50].

To study the contamination of the photon candidate spectrum, approximately  $3\times 10^6~\bar{n}$ 's were generated with an exponentially falling  $p_T$  spectrum. This provided sufficient statistics at low  $p_T$ , where the  $\bar{n}$ 's constituted a significant source of contamination. The FLUKA program [58] was used to describe the particle transport and the interactions in the detector material. The parametrizations of the p and  $\bar{p}$  yields were not only used to assign a weight to the Monte Carlo events, but also to determine the absolute contamination of the photon sample. The latter was divided by the number of reconstructed photon candidates to calculate the term  $C_0$ .

The final contamination factor  $C_0$  in the p + p data is shown in Fig. 17. In a limited  $p_T$  range,  $C_0$  appears to be larger than unity, which is not possible unless the associated systematic uncertainties are extremely large. The violation of this bound, as well as observed discrepancies between the three triggers, indicate the large uncertainty of the  $\bar{n}$  simulation.

Two natural limits to the contamination were considered: (i) the hard upper limit  $C_0 \le 1$ , which is not realistic because it implies that the inclusive photon yield is zero, and (ii) the limit implied by the assumption that the direct photon signal is zero around the annihilation peak ( $\approx 2~\text{GeV}/c$ ). In both cases, a scaling factor for  $C_0$  was derived and subsequently applied to the HighTower-2 data, as shown in Fig. 17. For the further analysis, we have chosen the second estimate, calculated assuming that only background photons were detected in the range  $1 < p_T < 4~\text{GeV}/c$ , as the upper limit of the neutral hadron contamination  $C_0$ . This upper limit was found

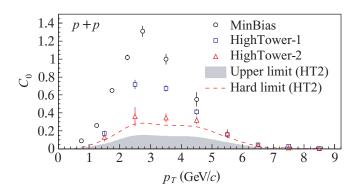


FIG. 17. (Color online) Relative neutral hadron contribution to the photon candidate yields,  $C_0$ . The filled area is the upper limit for the p + p HighTower-2 data sample; the dashed curve is the upper limit in the extreme scenario where all photon candidates were a result of showering  $\bar{n}$ 's.

to be negligible in the range of the present direct photon measurement,  $p_T > 6 \text{ GeV}/c$ .

The *n* and  $K_L^0$  interactions with the BEMC resulted in the smaller contamination than that of the  $\bar{n}$ 's at all values of  $p_T$ .

#### C. Photon reconstruction efficiency

We have calculated the photon acceptance and efficiency correction factor  $\varepsilon_{\rm acc}^{\gamma}$  separately for events containing  $\pi^0$  decay photons and for events containing only a single photon. The latter factor was applied to the fraction of the photon yield from all sources other than the  $\pi^0$  decay.

To determine the acceptance correction  $\varepsilon_{\rm acc}^{\gamma}$  for the  $\pi^0$  decay photons, we used a GEANT-based Monte Carlo simulation of the STAR detector. The Monte Carlo events were weighted in such a way that the measured  $\pi^0$  yield was reproduced. This is important because the photon acceptance depends on the degree of cluster merging. This, in turn, depends on the opening angle of the decay photons and thus on the momentum of the parent  $\pi^0$ . Furthermore, the simulation included all the possible losses of photon candidates listed in Sec. IV, except those associated with the invariant mass window and with the cut on the energy asymmetry  $Z_{\gamma\gamma}$ . One important effect affects the showers initiated by daughters of a high- $p_T$   $\pi^0$ . Because there was no requirement on the relation between the reconstructed  $p_T$  and the Monte Carlo input  $p_T$ , the correction implicitly accounted for events in which one of the two decay photons remained unidentified and the total energy was assigned to a single cluster. Such merging of photon showers constituted the main difference between the reconstruction efficiencies of  $\pi^0$  decay products and single

Similarly to the preceding, the  $\varepsilon_{\rm acc}^{\gamma}$  factor for single photons was determined using a Monte Carlo sample of  $\approx 1 \times 10^6$  events. Each event contained a single photon, uniformly distributed in azimuthal angle  $-\pi < \varphi < +\pi$ , pseudorapidity  $-0.3 < \eta < +1.2$ , and transverse momentum  $0 < p_T < 20~{\rm GeV}/c$ . Events were weighted with a function determined from the spectrum of photons from decaying hadrons other than the  $\pi^0$ , as well as from that of the direct photons. However, we demonstrate in Sec. VE that the shape of the decay photon spectrum and the measured  $\pi^0$  spectrum were very similar, at least for the  $p_T$  range of this analysis. Although the direct photon spectrum was expected to exhibit a slightly different  $p_T$  dependence, varying the input spectrum correspondingly did not yield quantitatively different results.

Finally, we have implemented a correction to the measured photon yields as follows. The yield of photons that originated from the decay  $\pi^0 \to \gamma \gamma$  was determined from the measured  $\pi^0$  spectrum. This part of  $Y_{\rm incl}$  was corrected with the  $\varepsilon_{\rm acc}^{\gamma}$  factor calculated for the  $\pi^0$  decay photons. The remaining part of  $Y_{\rm incl}$  was assumed to consist of single photons that were not correlated with the other photon candidates in the event and was corrected with the single-photon  $\varepsilon_{\rm acc}^{\gamma}$  factor. This assumption was based on the observation that the reconstruction efficiency for photons from the decay  $\eta \to \gamma \gamma$ , which is the second-largest source of decay photons ( $\approx$ 15%), was equal to that of single photons, because the opening angle

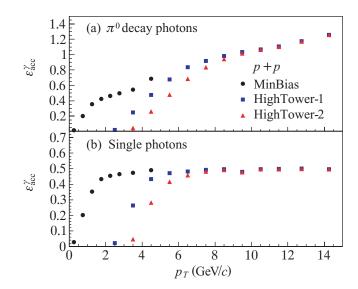


FIG. 18. (Color online) (a) Acceptance and efficiency factor  $\varepsilon_{\rm acc}^{\gamma}$  for photons exclusively from the decay  $\pi^0 \to \gamma \gamma$ . The observed rise of the efficiency at high  $p_T$  is caused by the merging of the  $\pi^0$  decay daughters. (b) The  $\varepsilon_{\rm acc}^{\gamma}$  factor for single photons, used to correct the fraction of the photon sample that exceeded the simulated  $\pi^0$  decay contribution.

between the two daughter photons is large enough that both are never incident on the same calorimeter tower.

Figure 18 shows the acceptance and reconstruction efficiency  $\varepsilon_{\rm acc}^{\gamma}$  for the  $\pi^0$  decay photons and for single photons for the p+p MinBias, HighTower-1, and HighTower-2 data. The two HighTower results were found to be very similar in the low- $p_T$  region, where the angular separation of the decay photons was still large, compared to the size of a BEMC tower. However, at higher  $p_T$  the two photons are difficult to separate, particularly in the cases of the most symmetric decays. When two photons were merged, the remaining photon candidate was erroneously assigned the energy sum of both showers. This led to significantly larger reconstruction efficiency, compared to that for single photons. Eventually, at the largest  $p_T$  values considered in this analysis, the decay photon efficiency even exceeded unity.

#### D. Fully corrected inclusive yields

The inclusive photon yield  $Y_{\rm incl}$  was derived by subtracting the charged and neutral backgrounds from the yield of raw photon candidates [Eq. (18)]. The contamination by charged particles was subtracted according to the procedure explained in Sec. III G, but the neutral hadron correction proved to be difficult. Although an upper limit for the contamination fraction  $C_0$  was derived in Sec. VB, we did not find any means of reducing the associated systematic uncertainty on  $C_0$  to a level where a meaningful subtraction could be performed for MinBias data. In the HighTower data, the upper limit on the contamination fraction vanishes at higher values of  $p_T$ . Therefore, our final results were obtained in the range  $6 < p_T < 15 \, {\rm GeV}/c$ , and the photon candidates obtained from the MinBias data were discarded.

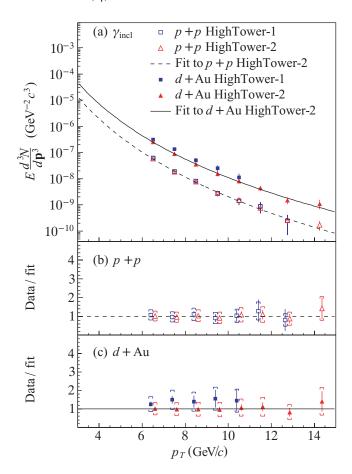


FIG. 19. (Color online) (a) Inclusive photon invariant yield per MinBias p+p and d+ Au collision. Curves are the power law fits given in the text. Invariant yield divided by the fit to the (b) p+p and (c) d+ Au data. The neutral hadron contamination was not subtracted (see text) and is expected to be significant at  $2 \lesssim p_T \lesssim 4 \text{ GeV}/c$ . The error bars are statistical and brackets in the lower panels are the systematic uncertainties. Data points for the overlapping  $p_T$  bins in the lower panels are horizontally displaced for clarity.

Figure 19 shows the corrected inclusive photon spectra in p+p and d+Au collisions without the subtraction of the neutral hadron contribution. Panels (b) and (c) show the data divided by the corresponding power law fits. A small systematic difference between the spectra from d+Au HighTower-1 and HighTower-2 collisions (1.2 standard deviations, on average) was observed. However, the measured HighTower-2 yields is statistically more significant, because the HighTower-1 events that contained photons in the range  $6 < p_T < 10 \text{ GeV}/c$  were a subset of the HighTower-2 data. We reconstructed 17684 (3738) photon candidates from the HighTower-2 (HighTower-1) d+Au data in that  $p_T$  range. The final direct photon cross sections presented in Sec. VI were obtained exclusively from the HighTower-2 data.

#### E. Background from hadronic decays

The photon yield from hadronic decays was determined with a simulation of the decay processes listed in Table IV. The other possible contributions, from processes such as  $\eta' \rightarrow$ 

TABLE IV. Dominant hadronic decay contributions to the inclusive photon yield [50].

Decay	Branching ratio (%)
$\pi^0  o \gamma \gamma$	98.80
$\pi^0  ightarrow e^+ e^- \gamma$	1.20
$\eta  o \gamma \gamma$	39.23
$\eta  o \pi^+\pi^-\gamma$	4.78
$\eta  ightarrow e^+ e^- \gamma$	0.49
$\omega(782) \to \pi^0 \gamma$	8.69

 $\rho^0 \gamma$ , were found to be negligible (<1%). A fit of the measured  $\pi^0$  yield in the range  $4 < p_T < 15~{\rm GeV}/c$  to the form  $\sim (1+p_T)^{-\alpha}$  served as an input to the simulation. The fit yielded  $\alpha = 9.1 \pm 0.1$  and  $9.0 \pm 0.1$  for  $d+{\rm Au}$  and p+p collisions, respectively. The normalization is irrelevant because it cancels in the ratio  $\gamma_{\rm decay}/\pi^0$ .

To estimate the yields of the  $\eta$  and  $\omega(782)$ , we used the fact that these scale with the  $\pi^0$  yields when expressed in terms of the transverse mass  $m_T \equiv \sqrt{m^2 + p_T^2}$  instead of  $p_T$  [25,59–61]. For the  $\eta$  spectra, we used the scaling ratios  $R_{\eta/\pi} = 0.46 \pm 0.05$  for p+p and  $0.44 \pm 0.08$  for d+Au data, as followed from our measurement of  $\eta$  production. In case of  $\omega(782)$ , we used  $R_{\omega/\pi} = 1.0 \pm 0.2$ , in agreement with recent measurements at RHIC [62]. The estimated  $\eta$  and  $\omega(782)$  yields relative to the  $\pi^0$  yield are shown in Fig. 20.

Figure 21 shows the ratio  $\gamma_{\rm decay}/\pi^0$  for  $d+{\rm Au}$  collisions. The curves represent the contributions of the  $\pi^0$ ,  $\eta$ , and  $\omega(782)$ , and the total decay photon yield, each divided by the parametrization of the measured  $\pi^0$  spectrum. The normalization uncertainty cancels upon taking this ratio. The uncertainty due to the shape of the  $\pi^0$  spectrum and to the  $m_T$  scaling factors was estimated by varying the fitted exponents and the scale factors by their errors.

## F. Summary of systematic uncertainties

All systematic error contributions to the double ratio  $R_{\gamma}$  [Eq. (15)] are summarized in Table V. Expressing the direct photon yield in terms of a double ratio gives a large reduction in the systematic error because the contribution from the BEMC energy scale uncertainty cancels. Consequently, the

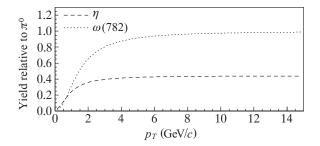


FIG. 20. The estimated yield of  $\eta$  and  $\omega(782)$  mesons in d+ Au collisions, relative to the measured  $\pi^0$  yield, determined from the  $m_T$  scaling as described in the text.

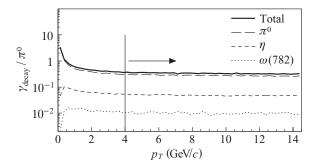


FIG. 21. The simulated number of photons per input pion  $\gamma_{\rm decay}/\pi^0$  from hadronic decays in  $d+{\rm Au}$  collisions as a function of  $p_T$ . The included decay processes are listed in Table IV. The vertical line indicates the lower limit of the  $p_T$  range that was used to fit the  $\pi^0$  spectrum. The value of  $\gamma_{\rm decay}/\pi^0$  below  $p_T=4~{\rm GeV}/c$  is, therefore, less accurate; however, it was not used for further analysis (see text).

largest sources of uncertainty are those associated with the  $\pi^0$  yield extraction and with the SMD energy scale. The latter leads to the  $R_{\gamma}$  variation of 12%, independent of  $p_T$  for  $p_T > 6 \, {\rm GeV}/c$ .

The beam background observed in the d+Au data has a larger effect on single-photon analysis than on the  $\pi^0$  reconstruction because the background-induced showers in the BEMC could not be distinguished from genuine photons originating from the event vertex. Therefore, we varied the cutoff value for the electromagnetic energy fraction r in an event [Eq. (1)] in the range r=0.7-0.9. This propagated into 1%-3% point-to-point systematic error of  $R_{\gamma}$ .

## VI. RESULTS AND DISCUSSION

# A. Cross section for neutral pion production

The invariant differential cross section for  $\pi^0$  and  $\eta$  production in inelastic p+p interactions is given by

$$E\frac{d^{3}\sigma_{\text{inel}}^{p+p}}{d\mathbf{p}^{3}} = E\frac{d^{3}\sigma_{\text{NSD}}^{p+p}}{d\mathbf{p}^{3}} = \sigma_{\text{NSD}}^{p+p}\frac{d^{2}N}{2\pi p_{T}dp_{T}dy}.$$
 (21)

It has been shown that the singly diffractive contribution to the inelastic cross section is negligible at  $p_T > 1 \text{ GeV}/c$  [23]. Therefore, we can assume that the differential inelastic cross

TABLE V. Systematic error contributions for the double ratio  $R_{\gamma}$ . The classifications A and B are defined in Sec. V.

Source	Type	Value at low (high) $p_T$ (%)
$\pi^0$ yield extraction	A	7.1
Beam background	A	1 (3) in $d + Au$
Tower energy scale	В	3
Tower gain spread	В	1
SMD energy scale	В	12
SMD gain spread	В	1
$\eta/\pi^0$	В	2
$\pi^0$ yield fit	В	1.5

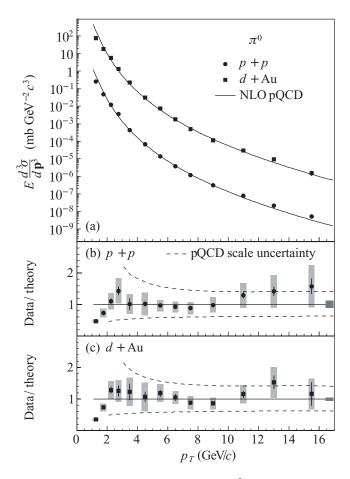


FIG. 22. (a) Cross section for inclusive  $\pi^0$  production in p+p and d+Au collisions at  $\sqrt{s_{NN}}=200$  GeV, divided by the corresponding NLO pQCD calculations [52] for (b) p+p and (c) d+Au collisions. The error bars are statistical and shaded bands are  $p_T$ -correlated systematic uncertainties. Normalization uncertainties are indicated by shaded bands around unity in the lower panels.

section is equal to the differential NSD cross section in our  $p_T$  range. The total NSD cross section in p+p collisions was found to be  $\sigma_{\rm NSD}^{p+p}=30.0\pm3.5$  mb, and the total hadronic cross section in  $d+{\rm Au}$  collisions was found to be  $\sigma_{\rm hadr}^{d+{\rm Au}}=2.21\pm0.09{\rm b}$  (see Sec. III A).

The measured cross sections for  $\pi^0$  production in the p+p(presented in Ref. [19] and included here for completeness) and d + Au collisions are shown in Fig. 22. The cross sections are compared to the NLO pOCD calculations [52]. The CTEO6M parton densities [63] and the KKP fragmentation functions [64] were used in the p + p calculation. The d + Au calculation used the nuclear parton distributions for gold [65–67], in addition. The factorization scale  $\mu$  was set equal to  $p_T$  and was varied by a factor of two to estimate the scale uncertainty, indicated by the dashed curves in Figs. 22(a) and 22(b). These panels show the ratio of the measured cross sections to the corresponding QCD predictions. The error bars shown in the plot are the statistical uncertainties and the shaded bands are the systematic uncertainties. The normalization uncertainties are indicated by shaded bands around unity on the right-hand side of each ratio plot. The measured  $\pi^0$  cross

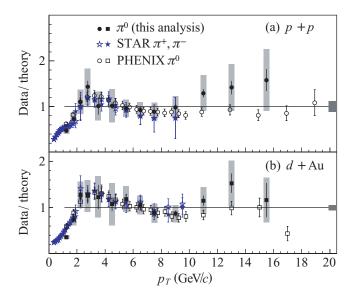


FIG. 23. (Color online) Cross section for inclusive  $\pi^0$  production in (a) p+p and (b) d+Au collisions at  $\sqrt{s_{NN}}=200$  GeV, divided by NLO pQCD calculations [52] and compared to the STAR  $\pi^{\pm}$  [23,24] and PHENIX  $\pi^0$  [6,26] measurements. The error bars are statistical and the shaded bands are  $p_T$ -correlated systematic uncertainties. Normalization uncertainties are indicated by shaded bands around unity in each panel.

sections were not corrected for feed-down contributions from  $\eta \to 3\pi^0, \eta \to \pi^+\pi^-\pi^0$ , and  $K_S^0 \to \pi^0\pi^0$  decays, which are expected to be negligible. It is seen that the measured  $\pi^0$  cross sections in both p+p and d+ Au collisions are well described by the NLO pQCD calculations in the fragmentation region  $p_T > 2~{\rm GeV}/c$ .

In Fig. 23, we compare the  $\pi^0$  measurements in the p+p and  $d+{\rm Au}$  data with the previous  $\pi^\pm$  measurements by STAR [23,24] and with the  $\pi^0$  measurements by PHENIX [6,26]. Here, and in all following figures, the cited data are shown with their statistical and systematic uncertainties added in quadrature. The normalization uncertainties shown by the gray bands in the figure are largely correlated between the  $\pi^0$  and the  $\pi^\pm$  data points and uncorrelated with the PHENIX normalization uncertainties of similar magnitude. It is seen that the neutral and charged pion spectra from STAR agree very well in both p+p and  $d+{\rm Au}$  data, in spite of different detector subsystems and analysis techniques used in these measurements. The present results extend the reach of STAR pion measurements to  $p_T=17~{\rm GeV}/c$ . Comparison to the

cross sections measured by PHENIX shows good agreement, within errors, in both collision systems. However, we note that our data indicate a possible excess over the PHENIX measurements at  $p_T > 10 \text{ GeV}/c$  in both cases.

To parametrize the  $p_T$  dependence, the measured  $\pi^0$  cross section, as well as the  $\eta$  and  $\gamma_{\rm incl}$  cross sections presented in the following sections, were fitted to the power law function [Eq. (13)], and the resulting parameter values are listed in Table VI. Because of the large uncertainties, the  $p_0$  and n parameters for the  $\eta$  cross section in d+ Au data had to be fixed at the corresponding p+p values to achieve a stable fit. The quoted values of  $\chi^2/$ ndf indicate that these fits provide only a general guidance on the shapes of the spectra and do not necessarily describe all features seen in the data.

In addition, the pure power law fit  $\sim p_T^{-m}$  to the  $\pi^0$  spectra at  $p_T > 5$  GeV/c gives  $m = 7.5 \pm 0.1$  ( $\chi^2$ /ndf = 6/5) for p + p and  $m = 7.9 \pm 0.2$  ( $\chi^2$ /ndf = 12/5) for d + Au collisions.

# B. $\eta$ -to-pion ratio

The  $\eta$  measurement is presented in Fig. 24 as the ratio of  $\eta$  to  $\pi^0$  invariant yields (shown in Figs. 16 and 15, respectively). This allows many systematic uncertainties to cancel (see Table III). The error definitions in the plot are the same as described previously for the differential cross sections. The present measurement agrees very well with previous PHENIX results (open symbols) [27]. The solid lines show the asymptotic ratio  $R^\infty=0.5$ , consistent with the world  $\eta/\pi^0$  measurements (see Ref. [27] and references therein). The fit to our data for  $p_T>4$  GeV/c gives  $R_{\eta/\pi}=0.46\pm0.05$  (p+p) and  $R_{\eta/\pi}=0.44\pm0.08$  (d+Au). The dashed curves in Fig. 24 show the prediction based on  $m_T$  scaling [25,59–61]. It is seen that the data are consistent with such scaling behavior.

# C. Nuclear modification factor

A convenient way to observe medium-induced modification of particle production is to compare a nucleus-nucleus collision (A+B) with an incoherent superposition of the corresponding number of individual nucleon-nucleon collisions (N+N). The nuclear modification factor  $R_{AB}$  is defined as the ratio of the particle yield in nucleus-nucleus collisions and the yield in nucleon-nucleon collisions scaled with the number of binary

TABLE VI. The values of the power law fit parameters from Eq. (13) for the measured  $\pi^0$ ,  $\eta$ , and  $\gamma_{\text{incl}}$  cross sections.

Data	$A \text{ (mb GeV}^{-2} c^3)$	$p_0  (\mathrm{GeV}/c)$	n	$\chi^2/\text{ndf}$
$\pi^0, p+p$	$(1.69 \pm 0.65) \times 10^3$	$0.723 \pm 0.066$	$8.61 \pm 0.14$	65/10
$\pi^0$ , $d + Au$	$(4.02 \pm 1.35) \times 10^4$	$1.46 \pm 0.16$	$9.93 \pm 0.32$	53/10
$\eta, p + p$	$(7.0 \pm 5.0) \times 10^{1}$	$1.33 \pm 0.23$	$9.83 \pm 0.44$	30/9
$\eta, d + Au$	$(3.33 \pm 0.41) \times 10^4$	1.33 (fixed)	9.83 (fixed)	32/10
$\gamma_{\rm incl}, p+p$	$(3.1 \pm 0.1) \times 10^{0}$	$0.941 \pm 0.268$	$8.61 \pm 0.40$	2/5
$\gamma_{\rm incl}$ , $d + Au$	$(2.4 \pm 0.1) \times 10^{1}$	$0.697 \pm 0.126$	$7.88 \pm 0.23$	2/5

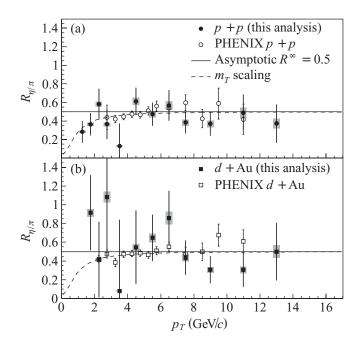


FIG. 24. The  $\eta/\pi^0$  ratio measured in (a) p+p and (b) d+1 Au collisions at  $\sqrt{s_{NN}} = 200$  GeV, compared to the PHENIX measurements [27] and the  $m_T$  scaling predictions. The error bars are statistical and the shaded bands are  $p_T$ -correlated systematic uncertainties.

collisions  $N_{\rm coll}$ ,

$$R_{AB} \equiv \frac{d^2 N_{AB}/dp_T dy}{\langle T_{AB} \rangle d^2 \sigma^{p+p}/dp_T dy}.$$
 (22)

Here  $\langle T_{AB} \rangle$  is the nuclear overlap function, which is related to the number of inelastic N+N collisions in one A+B collision through

$$\langle T_{AB} \rangle \sigma_{\text{inel}}^{NN} = \langle N_{\text{coll}} \rangle.$$
 (23)

In the absence of medium effects, the nuclear modification factor is unity, whereas  $R_{AB} < 1$  indicates a suppression of particle production in heavy-ion collisions, compared to an incoherent sum of nucleon-nucleon collisions.

We calculated the  $R_{dA}$  ratio [Eqs. (22) and (23)] as

$$R_{dA} = \frac{\sigma_{\text{inel}}^{NN} d^2 N_{dA} / dp_T dy}{\langle N_{\text{coll}} \rangle d^2 \sigma^{p+p} / dp_T dy},$$
(24)

where the nucleon-nucleon inelastic cross section was taken to be  $\sigma_{\rm inel}^{NN}=42$  mb and  $\langle N_{\rm coll}\rangle=7.5\pm0.4$  was calculated from the Glauber model (see Sec. III C).

The nuclear modification factors for  $\pi^0$  and  $\eta$  are shown in Fig. 25. The definition of the errors is the same as given for the differential cross sections in Sec. IV J. Figure 25 also shows the  $R_{dA}$  for  $\pi^\pm$  measured by STAR [23,24]. A good agreement between neutral and charged pion measurements by STAR is observed. Our  $\pi^0$  and  $\eta$  data also agree reasonably well with the corresponding PHENIX measurements [27,28]. In peripheral d+Au collisions, the number of participant nucleons is small and the creation of a dense medium is not expected. This suggests that, instead of p+p interactions, peripheral collisions can be used as a reference. This was done

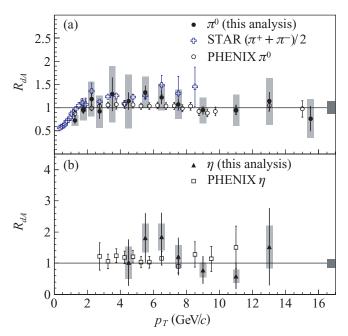


FIG. 25. (Color online) Nuclear modification factor  $R_{dA}$  for (a)  $\pi^0$  and (b)  $\eta$ , compared to the STAR  $\pi^{\pm}$  [23,24] and PHENIX  $\pi^0$  measurements [27,28]. The error bars are statistical and the shaded bands are  $p_T$ -correlated systematic uncertainties. Normalization uncertainties are indicated by shaded bands around unity in each panel.

through the ratio of particle production in 0%–20% central (C) and 40%–100% peripheral (P) events,

$$R_{CP} = \frac{\langle N_{\text{coll}} \rangle_P}{\langle N_{\text{coll}} \rangle_C} \frac{d^2 N_C / dp_T dy}{d^2 N_P / dp_T dy}.$$
 (25)

The advantage of this measure is that no p+p reference data are needed. The disadvantage is that a stronger model dependence is introduced due to the uncertainty in  $\langle N_{\rm coll} \rangle$ . Figure 26 shows the  $R_{CP}$  ratio for  $\pi^0$  compared to the STAR  $\pi^\pm$  data [24]. It is seen that the agreement between the neutral and the charged pion measurements is very good. The ratio stays constant at a value consistent with unity beyond  $p_T=8~{\rm GeV}/c$  and, therefore, does not support a possible decrease of the ratio at high  $p_T$ , which was suggested by the  $\pi^\pm$  measurement.

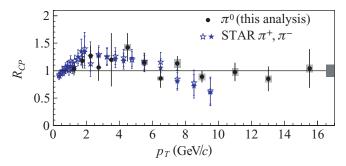


FIG. 26. (Color online) Nuclear modification factor  $R_{CP}$  measured in d + Au collisions, compared to STAR  $\pi^{\pm}$  measurement [24]. The error bars are statistical and the shaded bands are  $p_T$ -correlated systematic uncertainties. Common normalization uncertainty is indicated by a shaded band around unity.

## D. Direct photons

The double ratio  $R_{\gamma}$  [Eq. (15)] measured in p+p and d+ Au collisions is shown in Fig. 27. The shaded band near  $R_{\gamma}=0$  indicates our estimate of the upper limit of the remaining neutral hadron contamination. The curves correspond to NLO pQCD calculations [68], which were further evaluated as

$$R_{\gamma}|_{\text{theor}} = 1 + \frac{(\gamma_{\text{dir}}/\pi^0)_{\text{NLO}}}{(\gamma_{\text{decay}}/\pi^0)_{\text{simu}}},$$
 (26)

where the numerator is the ratio of the NLO pQCD direct photon and  $\pi^0$  cross sections. The denominator is given by the number of decay photons per  $\pi^0$ , as determined by the simulation described in Sec. V E.

The NLO pQCD calculation used the CTEQ6M [63] parton densities and the GRV [69] parton-to-photon fragmentation functions as an input. The scale dependence of this calculation, indicated by the dashed curves in the figure, was obtained by changing the scale  $\mu$  in the calculation of prompt photon production, while keeping the scale corresponding to the  $\pi^0$  cross section fixed at  $\mu=p_T$ . In addition, we have varied the factorization scale for both cross sections simultaneously. The observed variation was quantitatively similar, although in the opposite direction. Because the measured  $\pi^0$  spectrum favors the result of the pQCD calculation with  $\mu=p_T$ , we have used this value for all three curves.

Although Fig. 27 demonstrates that the measured values of  $R_{\gamma}$  are consistent with the calculated direct photon signal, the interpretation in this context has its limitations. First, the curves do not follow directly from the theory but depend on our simulation of the decay photon yields, as shown by Eq. (26). In addition, the NLO pQCD cross section for  $\pi^0$  production is less accurately constrained than that for prompt photon production.

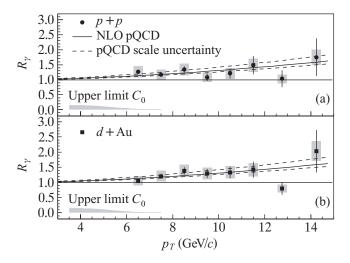


FIG. 27. The direct photon yield in (a) p+p and (b)  $d+{\rm Au}$  collisions at  $\sqrt{s_{NN}}=200$  GeV, expressed in terms of the double ratio  $R_{\gamma}$ . The error bars are statistical and the shaded bands are  $p_T$ -correlated systematic uncertainties. The curves correspond to NLO pQCD calculations of the differential cross sections for direct photon [68] and  $\pi^0$  [52] production in p+p collisions for different factorization scales  $\mu$  (the upper pQCD curve corresponds to  $\mu=p_T/2$ ). The upper limit of the fractional neutral hadron contamination  $C_0$  is shown as the shaded band at  $R_{\gamma}=0$ .

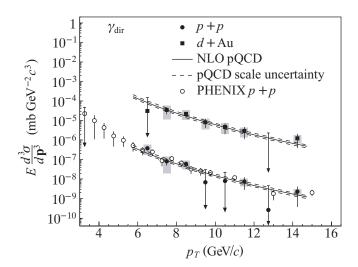


FIG. 28. The differential cross section for direct photon production at midrapidity in p+p and d+Au collisions at  $\sqrt{s_{NN}}=200$  GeV, compared to the PHENIX measurement [29] and the NLO pQCD calculation [68], which was scaled with  $\langle T_{dA} \rangle$  [Eq. (23)] in the case of d+Au collisions. The error bars are statistical and the the shaded bands are  $p_T$ -correlated systematic uncertainties. The arrows correspond to the 95% confidence limits, as defined in the text.

To allow for a more solid comparison to theoretical predictions, as well as to other experimental data, we have converted  $R_{\gamma}$  to an absolute cross section [Eq. (17)].

The calculation of absolute direct photon yields required that the systematic errors associated with inclusive photon yields, which canceled in the ratio  $R_{\gamma}$ , were included again. We derived the 95% confidence limits for the cross section in the  $p_T$  bins where  $R_{\gamma}$  did not correspond to a significant direct photon signal, assuming that the statistical and systematic errors both followed a Gaussian distribution and using the fact that  $R_{\gamma} \geqslant 1$  by definition.

Figure 28 shows the invariant cross section for direct photon production in p+p and d+Au collisions. The normalization uncertainties are not explicitly given in the figure. The NLO pQCD cross section for direct photon production in p+p collisions was scaled with the nuclear thickness function  $\langle T_{dA} \rangle$  [Eq. (23)] to account for the number of binary collisions in the d+Au system. The precision of the presented measurement is limited by systematic uncertainties for  $p_T \leq 9 \, \text{GeV}/c$  and by statistical uncertainties for larger  $p_T$  values. Nevertheless, our results are compatible with the NLO pQCD calculations. Our data are also in a good agreement with the direct photon cross section in p+p collisions measured by PHENIX [29].

Earlier measurements of direct photon production in proton-nucleus collisions have been performed by the E706 experiment [70] by scattering protons on a fixed beryllium target, with proton beam energies of 530 and 800 GeV. Those data show a strong discrepancy with pQCD calculations, which was attributed to multiple soft gluon radiation and phenomenologically described as an additional transverse impulse  $k_T$  to the incoming partons [70]. It has also been argued that this discrepancy might be attributable to nuclear modifications present even in the light berillium nucleus used

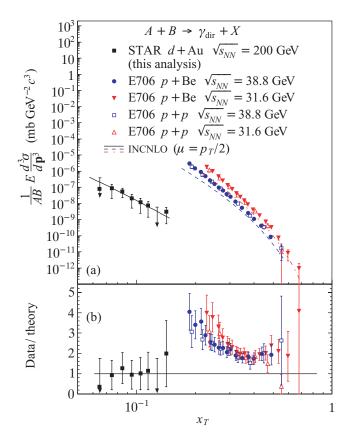


FIG. 29. (Color online) (a) The cross sections per nucleon for direct photon production in  $d+\mathrm{Au}$  collisions, compared to the measurements in p+p and p+A collisions by the E706 experiment [70] at the comparable nucleon-nucleon center-of-mass energies. The theoretical curves were calculated with the INCNLO program [72] with  $\mu=p_T/2$ . All data have statistical and systematic uncertainties summed in quadrature. The vertical arrows indicate our estimate of the 95% confidence level. (b) The ratio of the data and the corresponding calculations.

[71], although the p+p data by E706 show similar behavior at the same  $\sqrt{s_{NN}}$ . It is, therefore, of interest to compare our d+ Au results to those of E706, as shown in Fig. 29 as a function of  $x_T \equiv 2p_T/\sqrt{s}$ , which is a suitable variable for comparing data taken with different beam energies. Whereas the ratio data/theory from Ref. [70] shows an increase of up to a factor of 4 toward low  $x_T$ , our results at still lower  $x_T$  constrain such a potential deviation from theory to less than a factor of 2. It should be noted, however, that the data have been taken at significantly different  $\sqrt{s}$ .

We have included both prompt and fragmentation components in the pQCD calculations because our measurement was based on an inclusive sample of photons. The theoretical calculation of these two components is shown in Fig. 30. A first measurement of the contribution from fragmentation photons to the total direct photon cross section in p + p collisions at RHIC was reported in Ref. [29].

The interest in disentangling photons from the fragmentation process and from the initial hard scattering is twofold. First, it has been observed that the hot and dense medium

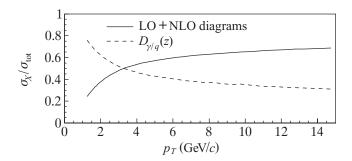


FIG. 30. The relative contribution of the fragmentation  $(D_{\gamma/q})$  and the pQCD hard scattering processes to the total direct photon cross section [68]. Both contributions are shown as a function of the  $p_T$  of the produced photon.

produced in central heavy-ion collisions at RHIC causes a suppression of particle yields, which has been attributed to induced gluon radiation from a parton traversing the medium. The same mechanism could lead to a suppression of fragmentation photons, although an enhancement of directly produced photons has been proposed as well [73]. Second, a measurement of the identified prompt photons in p+p collisions is of interest for the RHIC spin program, a large part of which is devoted to constraining the gluon spin contribution to the spin of proton. The isolation criterion selects the quark-gluon Compton process and, therefore, enhances the sensitivity of the cross section to the gluon content of the proton.

# E. Summary

The present  $\pi^0$  spectrum complements that of the  $\pi^{\pm}$ , which was measured by STAR in the transverse momentum range  $0.35 < p_T < 10 \text{ GeV}/c$ , and extends up to  $p_T = 17 \text{ GeV}/c$ . There is a good agreement between the neutral and the charged pion cross sections in STAR, even though very different methods and detector subsystems were used. The  $\pi^0$  cross section also agrees well with the measurements of PHENIX and with the theoretical NLO pQCD calculations.

This article presents the first measurements of  $\eta$  meson production by STAR, which are in agreement with the PHENIX measurements and with the  $m_T$  scaling assumption.

We present the measurements of the nuclear modification factor  $R_{dA}$ , where the  $\pi^0$  production in d+Au collisions is compared to that in p+p, and  $R_{CP}$ , the comparison between central and peripheral d+Au collisions. Both results are consistent with unity at high  $p_T$ , are in a good agreement with the  $\pi^\pm$  measurements, previously made by STAR, and significantly extend the  $p_T$  range for light meson production measurements.

This article also reports the first measurement of direct photon production by STAR. A direct photon signal consistent with NLO pQCD calculation has been observed at high  $p_T$  for both systems. No strong modification of photon production in d+ Au collisions was observed.

The results will provide an important baseline for future Au + Au measurements in STAR.

#### ACKNOWLEDGMENTS

We thank the RHIC Operations Group and RCF at BNL, the NERSC Center at LBNL and the Open Science Grid consortium for providing resources and support. This work was supported in part by the Offices of NP and HEP within the US DOE Office of Science, the US NSF, the Sloan Foundation, the DFG cluster of excellence "Origin and Structure of the Universe"; CNRS/IN2P3, STFC and EPSRC of the United

Kingdom; FAPESP CNPq of Brazil; Ministry of Education and Science of the Russian Federation; NNSFC, CAS, MoST, and MoE of China; GA and MSMT of the Czech Republic; FOM and NWO of the Netherlands; DAE, DST, and CSIR of India; the Polish Ministry of Science and Higher Education; Korea Research Foundation; Ministry of Science, Education and Sports of the Republic of Croatia; Russian Ministry of Science and Tech and RosAtom of Russia.

- [1] J. Adams *et al.* (STAR Collaboration), Phys. Rev. Lett. **91**, 172302 (2003).
- [2] C. Adler *et al.* (STAR Collaboration), Phys. Rev. Lett. 90, 082302 (2003).
- [3] C. Adler *et al.* (STAR Collaboration), Phys. Rev. Lett. 90, 032301 (2003).
- [4] J. Adams et al. (STAR Collaboration), Phys. Rev. Lett. 93, 252301 (2004).
- [5] J. Adams *et al.* (STAR Collaboration), Phys. Rev. Lett. 91, 072304 (2003).
- [6] S. Adler et al. (PHENIX Collaboration), Phys. Rev. Lett. 91, 072303 (2003).
- [7] B. B. Back *et al.* (PHOBOS Collaboration), Phys. Rev. Lett. 91, 072302 (2003).
- [8] I. Arsene *et al.* (BRAHMS Collaboration), Phys. Rev. Lett. 91, 072305 (2003).
- [9] M. Gyulassy, I. Vitev, X.-N. Wang, and B.-W. Zhang, in *Quark-Gluon Plasma*, edited by R. C. Hwa and X.-N. Wang (World Scientific Publishing Company Inc., 2004), p. 123.
- [10] J. Cronin et al., Phys. Rev. D 11, 3105 (1975).
- [11] S. Y. Li and X.-N. Wang, Phys. Lett. B 527, 85 (2002).
- [12] K. J. Eskola, V. J. Kolhinen, and C. A. Salgado, Eur. Phys. J. C 9, 61 (1999).
- [13] D. de Florian and R. Sassot, Phys. Rev. D 69, 074028 (2004).
- [14] B. Abelev et al. (STAR Collaboration), Phys. Rev. C 76, 054903 (2007).
- [15] X.-N. Wang, Z. Huang, and I. Sarcevic, Phys. Rev. Lett. 77, 231 (1996).
- [16] R. J. Fries, B. Müller, and D. K. Srivastava, Phys. Rev. Lett. 90, 132301 (2003).
- [17] T. Peitzmann and M. H. Thoma, Phys. Rep. **364**, 175 (2002)
- [18] G. Bunce, N. Saito, J. Soffer, and W. Vogelsang, Annu. Rev. Nucl. Part. Sci. 50, 525 (2000).
- [19] B. Abelev et al. (STAR Collaboration), Phys. Rev. D 80, 111108 (2009).
- [20] J. Adams et al. (STAR Collaboration), Phys. Rev. C 70, 044902 (2004).
- [21] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. C 80, 044905 (2009).
- [22] J. Adams et al. (STAR Collaboration), Phys. Rev. Lett. 97, 152302 (2006).
- [23] J. Adams et al. (STAR Collaboration), Phys. Lett. B 616, 8 (2005).
- [24] J. Adams et al. (STAR Collaboration), Phys. Lett. B 637, 161 (2006).
- [25] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. C **78**, 044906 (2008).
- [26] A. Adare *et al.* (PHENIX Collaboration), Phys. Rev. D **76**, 051106 (2007).

- [27] S. Adler et al. (PHENIX Collaboration), Phys. Rev. C 75, 024909 (2007).
- [28] S. Adler *et al.* (PHENIX Collaboration), Phys. Rev. Lett. 98, 172302 (2007).
- [29] S. Adler *et al.* (PHENIX Collaboration), Phys. Rev. Lett. 98, 012002 (2007).
- [30] K. Ackermann *et al.* (STAR Collaboration), Nucl. Instrum. Methods A **499**, 624 (2003).
- [31] M. Anderson et al. (STAR Collaboration), Nucl. Instrum. Methods A 499, 659 (2003).
- [32] K. Ackermann et al. (STAR Collaboration), Nucl. Instrum. Methods A 499, 713 (2003).
- [33] M. Beddo et al. (STAR Collaboration), Nucl. Instrum. Methods A 499, 725 (2003).
- [34] M. Bai et al., Nucl. Instrum. Methods A 499, 372 (2003).
- [35] C. Adler et al., Nucl. Instrum. Methods A 470, 488 (2001).
- [36] J. Kiryluk (STAR Collaboration), AIP Conf. Proc. 675, 424 (2003).
- [37] J. Kiryluk (STAR Collaboration), *16th International Spin Physics Symposium Proceedings* (World Scientific Publishing Company Inc., 2005), p. 718.
- [38] O. Grebenyuk, Ph.D. thesis, Utrecht University, 2007.
- [39] A. Drees and Z. Xu, *Proceedings of the Particle Accelerator Conference* (IEEE, Chicago, 2001), p. 3120.
- [40] M. L. Miller, K. Reygers, S. J. Sanders, and P. Steinberg, Annu. Rev. Nucl. Part. Sci. 57, 205 (2007).
- [41] C. Adler et al. (STAR Collaboration), Phys. Rev. Lett. 89, 202301 (2002).
- [42] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. C **79**, 034909 (2009).
- [43] A. Baltz et al., Nucl. Instrum. Methods A 417, 1 (1998).
- [44] L. Hulthén and M. Sugawara, in *Handbuch der Physik* (Springer-Verlag, Berlin, 1957), Vol. 39.
- [45] R. Brun, R. Hagelberg, M. Hansroul, and J. Lassalle, Technical Report CERN-DD-78-2-REV, CERN, 1978.
- [46] J. Adams et al. (STAR Collaboration), Phys. Rev. Lett. 92, 112301 (2004).
- [47] T. Cormier, A. Pavlinov, M. Rykov, V. Rykov, and K. Shestermanov, Nucl. Instrum. Methods A 483, 734 (2002).
- [48] A. Wetzler, Ph.D. thesis, J. W. Goethe-University, 2006.
- [49] M. Russcher, Ph.D. thesis, Utrecht University, 2008.
- [50] C. Amsler *et al.* (Particle Data Group), Phys. Lett. B 667, 1 (2008).
- [51] B. Abelev et al. (STAR Collaboration), Phys. Rev. Lett. 97, 252001 (2006).
- [52] B. Jäger, A. Schäfer, M. Stratmann, and W. Vogelsang, Phys. Rev. D 67, 054005 (2003).
- [53] T. Sjöstrand, P. Edén, C. Friberg, L. Lönnblad, G. Miu, S. Mrenna, and E. Norrbin, Comput. Phys. Commun. 135, 238 (2001).

- [54] G. D. Lafferty and T. R. Wyatt, Nucl. Instrum. Methods A 355, 541 (1995).
- [55] P. Aurenche, J. P. Guillet, E. Pilon, M. Werlen, and M. Fontannaz, Phys. Rev. D 73, 094007 (2006).
- [56] M. M. Aggarwal *et al.* (WA98 Collaboration), Phys. Rev. Lett. 85, 3595 (2000).
- [57] S. Adler *et al.* (PHENIX Collaboration), Phys. Rev. Lett. 94, 232301 (2005).
- [58] A. Fasso et al., Computing in High Energy and Nuclear Physics Conference Proceeding (La Jolla, CA, 2003), p. MOMT005.
- [59] M. Bourquin and J.-M. Gaillard, Nucl. Phys. B 114, 334 (1976).
- [60] J. Schaffner-Bielich, D. Kharzeev, L. McLerran, and R. Venugopalan, Nucl. Phys. A 705, 494 (2002).
- [61] B. I. Abelev et al. (STAR Collaboration), Phys. Rev. C 75, 064901 (2007).
- [62] S. Adler et al. (PHENIX Collaboration), Phys. Rev. C 75, 051902 (2007).
- [63] J. Pumplin et al., J. High Energy Phys. 07 (2002) 012.

- [64] B. Kniehl, G. Kramer, and B. Pötter, Nucl. Phys. B 582, 514 (2000).
- [65] L. Frankfurt and M. Strikman, Eur. Phys. J. A 5, 293 (1999).
- [66] L. Frankfurt, V. Guzey, M. McDermott, and M. Strikman, J. High Energy Phys. 02 (2002) 027.
- [67] L. Frankfurt, V. Guzey, and M. Strikman, Phys. Rev. D 71, 054001 (2005).
- [68] L. E. Gordon and W. Vogelsang, Phys. Rev. D 48, 3136 (1993).
- [69] M. Glück, E. Reya, and A. Vogt, Phys. Rev. D 48, 116 (1993).
- [70] L. Apanasevich *et al.* (E706 Collaboration), Phys. Rev. D 70, 092009 (2004).
- [71] G. Papp, P. Lévai, and G. Fai, Phys. Rev. C 61, 021902 (1999).
- [72] P. Aurenche, M. Fontannaz, J. P. Guillet, B. Kniehl, E. Pilon, and M. Werlen, Eur. Phys. J. C 9, 107 (1999).
- [73] R. J. Fries, B. Müller, and D. K. Srivastava, Phys. Rev. Lett. 90, 132301 (2003).