

# Precision Measurement of the Longitudinal Double-Spin Asymmetry for Inclusive Jet Production in Polarized Proton Collisions at $\sqrt{s}=200$ GeV

---

(STAR Collaboration) Adamczyk, L.; ...; Planinić, Mirko; ...; Poljak, Nikola; ...; Zyzak, M.

Source / Izvornik: **Physical Review Letters, 2015, 115**

Journal article, Published version

Rad u časopisu, Objavljena verzija rada (izdavačev PDF)

<https://doi.org/10.1103/PhysRevLett.115.092002>

Permanent link / Trajna poveznica: <https://um.nsk.hr/um:nbn:hr:217:148898>

Rights / Prava: [In copyright](#)/[Zaštićeno autorskim pravom](#).

Download date / Datum preuzimanja: **2025-04-01**



Repository / Repozitorij:

[Repository of the Faculty of Science - University of Zagreb](#)





## Precision Measurement of the Longitudinal Double-Spin Asymmetry for Inclusive Jet Production in Polarized Proton Collisions at $\sqrt{s} = 200$ GeV

L. Adamczyk,<sup>1</sup> J. K. Adkins,<sup>23</sup> G. Agakishiev,<sup>21</sup> M. M. Aggarwal,<sup>35</sup> Z. Ahammed,<sup>53</sup> I. Alekseev,<sup>19</sup> J. Alford,<sup>22</sup> C. D. Anson,<sup>32</sup> A. Aparin,<sup>21</sup> D. Arkhipkin,<sup>4</sup> E. C. Aschenauer,<sup>4</sup> G. S. Averichev,<sup>21</sup> A. Banerjee,<sup>53</sup> D. R. Beavis,<sup>4</sup> R. Bellwied,<sup>49</sup> A. Bhasin,<sup>20</sup> A. K. Bhati,<sup>35</sup> P. Bhattarai,<sup>48</sup> H. Bichsel,<sup>55</sup> J. Bielcik,<sup>13</sup> J. Bielcikova,<sup>14</sup> L. C. Bland,<sup>4</sup> I. G. Bordyuzhin,<sup>19</sup> W. Borowski,<sup>45</sup> J. Bouchet,<sup>22</sup> A. V. Brandin,<sup>30</sup> S. G. Brovko,<sup>6</sup> S. Bültmann,<sup>33</sup> I. Bunzarov,<sup>21</sup> T. P. Burton,<sup>4</sup> J. Butterworth,<sup>41</sup> H. Caines,<sup>57</sup> M. Calderón de la Barca Sánchez,<sup>6</sup> J. M. Campbell,<sup>32</sup> D. Cebra,<sup>6</sup> R. Cendejas,<sup>36</sup> M. C. Cervantes,<sup>47</sup> P. Chaloupka,<sup>13</sup> Z. Chang,<sup>47</sup> S. Chattopadhyay,<sup>53</sup> H. F. Chen,<sup>42</sup> J. H. Chen,<sup>44</sup> L. Chen,<sup>9</sup> J. Cheng,<sup>50</sup> M. Cherney,<sup>12</sup> A. Chikanian,<sup>57</sup> W. Christie,<sup>4</sup> J. Chwastowski,<sup>11</sup> M. J. M. Coddington,<sup>48</sup> G. Contin,<sup>26</sup> J. G. Cramer,<sup>55</sup> H. J. Crawford,<sup>5</sup> A. B. Cudd,<sup>47</sup> X. Cui,<sup>42</sup> S. Das,<sup>16</sup> A. Davila Leyva,<sup>48</sup> L. C. De Silva,<sup>12</sup> R. R. Debbe,<sup>4</sup> T. G. Dedovich,<sup>21</sup> J. Deng,<sup>43</sup> A. A. Derevschikov,<sup>37</sup> R. Derradi de Souza,<sup>8</sup> S. Dhamija,<sup>18</sup> B. di Ruzza,<sup>4</sup> L. Didenko,<sup>4</sup> C. Dilks,<sup>36</sup> F. Ding,<sup>6</sup> P. Djawotho,<sup>47</sup> X. Dong,<sup>26</sup> J. L. Drachenberg,<sup>52</sup> J. E. Draper,<sup>6</sup> C. M. Du,<sup>25</sup> L. E. Dunkelberger,<sup>7</sup> J. C. Dunlop,<sup>4</sup> L. G. Efimov,<sup>21</sup> J. Engelage,<sup>5</sup> K. S. Engle,<sup>51</sup> G. Eppley,<sup>41</sup> L. Eun,<sup>26</sup> O. Evdokimov,<sup>10</sup> O. Eyser,<sup>4</sup> R. Fatemi,<sup>23</sup> S. Fazio,<sup>4</sup> J. Fedorisin,<sup>21</sup> P. Filip,<sup>21</sup> E. Finch,<sup>57</sup> Y. Fisyak,<sup>4</sup> C. E. Flores,<sup>6</sup> C. A. Gagliardi,<sup>47</sup> D. R. Gangadharan,<sup>32</sup> D. Garand,<sup>38</sup> F. Geurts,<sup>41</sup> A. Gibson,<sup>52</sup> M. Girard,<sup>54</sup> S. Gliske,<sup>2</sup> L. Greiner,<sup>26</sup> D. Grosnick,<sup>52</sup> D. S. Gunarathne,<sup>46</sup> Y. Guo,<sup>42</sup> A. Gupta,<sup>20</sup> S. Gupta,<sup>20</sup> W. Guryn,<sup>4</sup> B. Haag,<sup>6</sup> A. Hamed,<sup>47</sup> L.-X. Han,<sup>44</sup> R. Haque,<sup>31</sup> J. W. Harris,<sup>57</sup> S. Heppelmann,<sup>36</sup> A. Hirsch,<sup>38</sup> G. W. Hoffmann,<sup>48</sup> D. J. Hofman,<sup>10</sup> S. Horvat,<sup>57</sup> B. Huang,<sup>4</sup> H. Z. Huang,<sup>7</sup> X. Huang,<sup>50</sup> P. Huck,<sup>9</sup> T. J. Humanic,<sup>32</sup> G. Igo,<sup>7</sup> W. W. Jacobs,<sup>18</sup> H. Jang,<sup>24</sup> E. G. Judd,<sup>5</sup> S. Kabana,<sup>45</sup> D. Kalinkin,<sup>19</sup> K. Kang,<sup>50</sup> K. Kauder,<sup>10</sup> H. W. Ke,<sup>4</sup> D. Keane,<sup>22</sup> A. Kechechyan,<sup>21</sup> A. Kesich,<sup>6</sup> Z. H. Khan,<sup>10</sup> D. P. Kikola,<sup>54</sup> I. Kisel,<sup>15</sup> A. Kisiel,<sup>54</sup> D. D. Koetke,<sup>52</sup> T. Kollegger,<sup>15</sup> J. Konzer,<sup>38</sup> I. Koralt,<sup>33</sup> L. K. Kosarzewski,<sup>54</sup> L. Kotchenda,<sup>30</sup> A. F. Kraishan,<sup>46</sup> P. Kravtsov,<sup>30</sup> K. Krueger,<sup>2</sup> I. Kulakov,<sup>15</sup> L. Kumar,<sup>31</sup> R. A. Kycia,<sup>11</sup> M. A. C. Lamont,<sup>4</sup> J. M. Landgraf,<sup>4</sup> K. D. Landry,<sup>7</sup> J. Lauret,<sup>4</sup> A. Lebedev,<sup>4</sup> R. Lednicky,<sup>21</sup> J. H. Lee,<sup>4</sup> M. J. LeVine,<sup>4</sup> C. Li,<sup>42</sup> W. Li,<sup>44</sup> X. Li,<sup>38</sup> X. Li,<sup>46</sup> Y. Li,<sup>50</sup> Z. M. Li,<sup>9</sup> M. A. Lisa,<sup>32</sup> F. Liu,<sup>9</sup> T. Ljubicic,<sup>4</sup> W. J. Llope,<sup>41</sup> M. Lomnitz,<sup>22</sup> R. S. Longacre,<sup>4</sup> X. Luo,<sup>9</sup> G. L. Ma,<sup>44</sup> Y. G. Ma,<sup>44</sup> D. M. M. D. Madagodagetige Don,<sup>12</sup> D. P. Mahapatra,<sup>16</sup> R. Majka,<sup>57</sup> S. Margetis,<sup>22</sup> C. Markert,<sup>48</sup> H. Masui,<sup>26</sup> H. S. Matis,<sup>26</sup> D. McDonald,<sup>49</sup> T. S. McShane,<sup>12</sup> N. G. Minaev,<sup>37</sup> S. Mioduszewski,<sup>47</sup> B. Mohanty,<sup>31</sup> M. M. Mondal,<sup>47</sup> D. A. Morozov,<sup>37</sup> M. K. Mustafa,<sup>26</sup> B. K. Nandi,<sup>17</sup> Md. Nasim,<sup>31</sup> T. K. Nayak,<sup>53</sup> J. M. Nelson,<sup>3</sup> G. Nigmatkulov,<sup>30</sup> L. V. Nogach,<sup>37</sup> S. Y. Noh,<sup>24</sup> J. Novak,<sup>29</sup> S. B. Nurushev,<sup>37</sup> G. Odyniec,<sup>26</sup> A. Ogawa,<sup>4</sup> K. Oh,<sup>39</sup> A. Ohlson,<sup>57</sup> V. Okorokov,<sup>30</sup> E. W. Oldag,<sup>48</sup> D. L. Olivitt, Jr.,<sup>46</sup> M. Pachr,<sup>13</sup> B. S. Page,<sup>18</sup> S. K. Pal,<sup>53</sup> Y. X. Pan,<sup>7</sup> Y. Pandit,<sup>10</sup> Y. Panebratsev,<sup>21</sup> T. Pawlak,<sup>54</sup> B. Pawlik,<sup>34</sup> H. Pei,<sup>9</sup> C. Perkins,<sup>5</sup> W. Peryt,<sup>54</sup> P. Pile,<sup>4</sup> M. Planinic,<sup>58</sup> J. Pluta,<sup>54</sup> N. Poljak,<sup>58</sup> K. Poniatowska,<sup>54</sup> J. Porter,<sup>26</sup> A. M. Poskanzer,<sup>26</sup> N. K. Pruthi,<sup>35</sup> M. Przybycien,<sup>1</sup> P. R. Pujahari,<sup>17</sup> J. Putschke,<sup>56</sup> H. Qiu,<sup>26</sup> A. Quintero,<sup>22</sup> S. Ramachandran,<sup>23</sup> R. Raniwala,<sup>40</sup> S. Raniwala,<sup>40</sup> R. L. Ray,<sup>48</sup> C. K. Riley,<sup>57</sup> H. G. Ritter,<sup>26</sup> J. B. Roberts,<sup>41</sup> O. V. Rogachevskiy,<sup>21</sup> J. L. Romero,<sup>6</sup> J. F. Ross,<sup>12</sup> A. Roy,<sup>53</sup> L. Ruan,<sup>4</sup> J. Rusnak,<sup>14</sup> O. Rusnakova,<sup>13</sup> N. R. Sahoo,<sup>47</sup> P. K. Sahu,<sup>16</sup> I. Sakrejda,<sup>26</sup> S. Salur,<sup>26</sup> J. Sandweiss,<sup>57</sup> E. Sangaline,<sup>6</sup> A. Sarkar,<sup>17</sup> J. Schambach,<sup>48</sup> R. P. Scharenberg,<sup>38</sup> A. M. Schmah,<sup>26</sup> W. B. Schmidke,<sup>4</sup> N. Schmitz,<sup>28</sup> J. Seger,<sup>12</sup> P. Seyboth,<sup>28</sup> N. Shah,<sup>7</sup> E. Shabaliev,<sup>21</sup> P. V. Shanmuganathan,<sup>22</sup> M. Shao,<sup>42</sup> B. Sharma,<sup>35</sup> W. Q. Shen,<sup>44</sup> S. S. Shi,<sup>26</sup> Q. Y. Shou,<sup>44</sup> E. P. Sichtermann,<sup>26</sup> R. N. Singaraju,<sup>53</sup> M. J. Skoby,<sup>18</sup> D. Smirnov,<sup>4</sup> N. Smirnov,<sup>57</sup> D. Solanki,<sup>40</sup> P. Sorensen,<sup>4</sup> H. M. Spinka,<sup>2</sup> B. Srivastava,<sup>38</sup> T. D. S. Stanislaus,<sup>52</sup> J. R. Stevens,<sup>27</sup> R. Stock,<sup>15</sup> M. Strikhanov,<sup>30</sup> B. Stringfellow,<sup>38</sup> M. Sumbera,<sup>14</sup> X. Sun,<sup>26</sup> X. M. Sun,<sup>26</sup> Y. Sun,<sup>42</sup> Z. Sun,<sup>25</sup> B. Surrow,<sup>46</sup> D. N. Svirida,<sup>19</sup> T. J. M. Symons,<sup>26</sup> M. A. Szelezniak,<sup>26</sup> J. Takahashi,<sup>8</sup> A. H. Tang,<sup>4</sup> Z. Tang,<sup>42</sup> T. Tarnowsky,<sup>29</sup> J. H. Thomas,<sup>26</sup> A. R. Timmins,<sup>49</sup> D. Tlusty,<sup>14</sup> M. Tokarev,<sup>21</sup> S. Trentalange,<sup>7</sup> R. E. Tribble,<sup>47</sup> P. Tribedy,<sup>53</sup> B. A. Trzeciak,<sup>13</sup> O. D. Tsai,<sup>7</sup> J. Turnau,<sup>34</sup> T. Ullrich,<sup>4</sup> D. G. Underwood,<sup>2</sup> G. Van Buren,<sup>4</sup> G. van Nieuwenhuizen,<sup>27</sup> M. Vandenbroucke,<sup>46</sup> J. A. Vanfossen, Jr.,<sup>22</sup> R. Varma,<sup>17</sup> G. M. S. Vasconcelos,<sup>8</sup> A. N. Vasiliev,<sup>37</sup> R. Vertesi,<sup>14</sup> F. Videbæk,<sup>4</sup> Y. P. Vijoyi,<sup>53</sup> S. Vokal,<sup>21</sup> A. Vossen,<sup>18</sup> M. Wada,<sup>48</sup> F. Wang,<sup>38</sup> G. Wang,<sup>7</sup> H. Wang,<sup>4</sup> J. S. Wang,<sup>25</sup> X. L. Wang,<sup>42</sup> Y. Wang,<sup>50</sup> Y. Wang,<sup>10</sup> G. Webb,<sup>23</sup> J. C. Webb,<sup>4</sup> G. D. Westfall,<sup>29</sup> H. Wieman,<sup>26</sup> S. W. Wissink,<sup>18</sup> R. Witt,<sup>51</sup> Y. F. Wu,<sup>9</sup> Z. Xiao,<sup>50</sup> W. Xie,<sup>38</sup> K. Xin,<sup>41</sup> H. Xu,<sup>25</sup> J. Xu,<sup>9</sup> N. Xu,<sup>26</sup> Q. H. Xu,<sup>43</sup> Y. Xu,<sup>42</sup> Z. Xu,<sup>4</sup> W. Yan,<sup>50</sup> C. Yang,<sup>42</sup> Y. Yang,<sup>25</sup> Y. Yang,<sup>9</sup> Z. Ye,<sup>10</sup> P. Yepes,<sup>41</sup> L. Yi,<sup>38</sup> K. Yip,<sup>4</sup> I.-K. Yoo,<sup>39</sup> N. Yu,<sup>9</sup> Y. Zawisza,<sup>42</sup> H. Zbroszczyk,<sup>54</sup> W. Zha,<sup>42</sup> J. B. Zhang,<sup>9</sup> J. L. Zhang,<sup>43</sup> S. Zhang,<sup>44</sup> X. P. Zhang,<sup>50</sup> Y. Zhang,<sup>42</sup> Z. P. Zhang,<sup>42</sup> F. Zhao,<sup>7</sup> J. Zhao,<sup>9</sup> C. Zhong,<sup>44</sup> X. Zhu,<sup>50</sup> Y. H. Zhu,<sup>44</sup> Y. Zoukarnieva,<sup>21</sup> and M. Zyzak<sup>15</sup>

(STAR Collaboration)

- <sup>1</sup>AGH University of Science and Technology, Cracow 30-059, Poland  
<sup>2</sup>Argonne National Laboratory, Argonne, Illinois 60439, USA  
<sup>3</sup>University of Birmingham, Birmingham, United Kingdom  
<sup>4</sup>Brookhaven National Laboratory, Upton, New York 11973, USA  
<sup>5</sup>University of California, Berkeley, California 94720, USA  
<sup>6</sup>University of California, Davis, California 95616, USA  
<sup>7</sup>University of California, Los Angeles, California 90095, USA  
<sup>8</sup>Universidade Estadual de Campinas, Sao Paulo 05314-970, Brazil  
<sup>9</sup>Central China Normal University (HZNU), Wuhan 430079, China  
<sup>10</sup>University of Illinois at Chicago, Chicago, Illinois 60607, USA  
<sup>11</sup>Cracow University of Technology, Cracow 31-342, Poland  
<sup>12</sup>Creighton University, Omaha, Nebraska 68178, USA  
<sup>13</sup>Czech Technical University in Prague, FNSPE, Prague 115 19, Czech Republic  
<sup>14</sup>Nuclear Physics Institute AS CR, 250 68 Řež/Prague, Czech Republic  
<sup>15</sup>Frankfurt Institute for Advanced Studies FIAS, Frankfurt am Main D-60438, Germany  
<sup>16</sup>Institute of Physics, Bhubaneswar 751005, India  
<sup>17</sup>Indian Institute of Technology, Mumbai 400076, India  
<sup>18</sup>Indiana University, Bloomington, Indiana 47408, USA  
<sup>19</sup>Alikhanov Institute for Theoretical and Experimental Physics, Moscow 117218, Russia  
<sup>20</sup>University of Jammu, Jammu 180001, India  
<sup>21</sup>Joint Institute for Nuclear Research, Dubna 141 980, Russia  
<sup>22</sup>Kent State University, Kent, Ohio 44242, USA  
<sup>23</sup>University of Kentucky, Lexington, Kentucky 40506-0055, USA  
<sup>24</sup>Korea Institute of Science and Technology Information, Daejeon 305-806, Korea  
<sup>25</sup>Institute of Modern Physics, Lanzhou 730000, China  
<sup>26</sup>Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA  
<sup>27</sup>Massachusetts Institute of Technology, Cambridge, Massachusetts 02139-4307, USA  
<sup>28</sup>Max-Planck-Institut für Physik, Munich D-80805, Germany  
<sup>29</sup>Michigan State University, East Lansing, Michigan 48824, USA  
<sup>30</sup>Moscow Engineering Physics Institute, Moscow 115409, Russia  
<sup>31</sup>National Institute of Science Education and Research, Bhubaneswar 751005, India  
<sup>32</sup>Ohio State University, Columbus, Ohio 43210, USA  
<sup>33</sup>Old Dominion University, Norfolk, Virginia 23529, USA  
<sup>34</sup>Institute of Nuclear Physics PAN, Cracow 31-342, Poland  
<sup>35</sup>Panjab University, Chandigarh 160014, India  
<sup>36</sup>Pennsylvania State University, University Park, Pennsylvania 16802, USA  
<sup>37</sup>Institute of High Energy Physics, Protvino 142281, Russia  
<sup>38</sup>Purdue University, West Lafayette, Indiana 47907, USA  
<sup>39</sup>Pusan National University, Pusan 609-735, Republic of Korea  
<sup>40</sup>University of Rajasthan, Jaipur 302004, India  
<sup>41</sup>Rice University, Houston, Texas 77251, USA  
<sup>42</sup>University of Science and Technology of China, Hefei 230026, China  
<sup>43</sup>Shandong University, Jinan, Shandong 250100, China  
<sup>44</sup>Shanghai Institute of Applied Physics, Shanghai 201800, China  
<sup>45</sup>SUBATECH, Nantes, France  
<sup>46</sup>Temple University, Philadelphia, Pennsylvania 19122, USA  
<sup>47</sup>Texas A&M University, College Station, Texas 77843, USA  
<sup>48</sup>University of Texas, Austin, Texas 78712, USA  
<sup>49</sup>University of Houston, Houston, Texas 77204, USA  
<sup>50</sup>Tsinghua University, Beijing 100084, China  
<sup>51</sup>United States Naval Academy, Annapolis, Maryland 21402, USA  
<sup>52</sup>Valparaiso University, Valparaiso, Indiana 46383, USA  
<sup>53</sup>Variable Energy Cyclotron Centre, Kolkata 700064, India  
<sup>54</sup>Warsaw University of Technology, Warsaw 00 662, Poland  
<sup>55</sup>University of Washington, Seattle, Washington 98195, USA  
<sup>56</sup>Wayne State University, Detroit, Michigan 48201, USA  
<sup>57</sup>Yale University, New Haven, Connecticut 06520, USA  
<sup>58</sup>University of Zagreb, Zagreb HR-10002, Croatia

(Received 20 May 2014; revised manuscript received 7 May 2015; published 26 August 2015)

We report a new measurement of the midrapidity inclusive jet longitudinal double-spin asymmetry,  $A_{LL}$ , in polarized  $pp$  collisions at center-of-mass energy  $\sqrt{s} = 200$  GeV. The STAR data place stringent constraints on polarized parton distribution functions extracted at next-to-leading order from global analyses of inclusive deep-inelastic scattering (DIS), semi-inclusive DIS, and RHIC  $pp$  data. The measured asymmetries provide evidence at the  $3\sigma$  level for positive gluon polarization in the Bjorken- $x$  region  $x > 0.05$ .

DOI: [10.1103/PhysRevLett.115.092002](https://doi.org/10.1103/PhysRevLett.115.092002)

PACS numbers: 14.20.Dh, 13.87.Ce, 13.88.+e, 14.70.Dj

A fundamental and long-standing puzzle in quantum chromodynamics (QCD) concerns how the intrinsic spins and orbital angular momenta of the quarks, antiquarks, and gluons sum to give the proton spin of  $\hbar/2$  [1]. The flavor-summed quark and antiquark spin contributions,  $\Delta\Sigma$ , account for less than a third of the total proton spin [2–6]. Because of the limited range in momentum transfer at a fixed Bjorken  $x$  accessed by fixed-target experiments, the polarized deep-inelastic scattering (DIS) data used to extract  $\Delta\Sigma$  provide only loose constraints on the gluon spin contribution,  $\Delta G$ , via scaling violations.

The measurement of asymmetries directly sensitive to the gluon helicity distribution was a primary motivation for establishing the spin structure program at the Relativistic Heavy Ion Collider (RHIC). Since the commencement of the RHIC spin program, several inclusive jet [7–9] and pion [10–14] asymmetry measurements have been incorporated into next-to-leading-order (NLO) perturbative QCD (pQCD) fits. While these data provide some constraints on  $\Delta G$ , ruling out large positive or negative gluon contributions to the proton spin, they lack the statistical power to distinguish a moderate gluon contribution from zero. The inclusive jet asymmetries presented here benefit from nearly a 20-fold increase in the event sample as well as improved jet reconstruction and correction techniques compared to Ref. [9], and they provide much tighter constraints on the gluon polarization.

The cross section for midrapidity inclusive jet production in  $pp$  collisions at  $\sqrt{s} = 200$  GeV is well described by NLO pQCD calculations [15,16] over the transverse momentum range  $5 < p_T < 50$  GeV/ $c$  [7]. The NLO pQCD calculations indicate that midrapidity jet production at RHIC is dominated by quark-gluon ( $qg$ ) and gluon-gluon ( $gg$ ) scattering, which together account for 60%–90% of the total yield for the jet transverse momenta studied here. The  $qg$  and  $gg$  scattering cross sections are very sensitive to the longitudinal helicities of the participating partons, so the inclusive jet longitudinal double-spin asymmetry,  $A_{LL}$ , provides direct sensitivity to the gluon polarization in the proton.  $A_{LL}$  is defined as

$$A_{LL} = \frac{\sigma^{++} - \sigma^{+-}}{\sigma^{++} + \sigma^{+-}}, \quad (1)$$

where  $\sigma^{++}(\sigma^{+-})$  is the differential cross section when the beam protons have the same (opposite) helicities.

The data presented here were extracted from an integrated luminosity of  $20 \text{ pb}^{-1}$  recorded in 2009 with the STAR detector [17] at RHIC. The polarization was measured independently for each of the two counterrotating proton beams [hereafter designated blue ( $B$ ) and yellow ( $Y$ )] and for each fill using Coulomb-nuclear interference proton-carbon polarimeters [18], calibrated via a polarized atomic hydrogen gas-jet target [19]. Averaged over RHIC fills, the luminosity-weighted polarization values for the two beams were  $P_B = 0.574$  and  $P_Y = 0.573$ , with a 6.5% relative uncertainty on the product  $P_B P_Y$  [20]. The helicity patterns of the colliding beam bunches were changed between beam fills to minimize systematic uncertainties in the  $A_{LL}$  measurement. The asymmetry  $A_{LL}$  is determined from the ratio of yields for different beam-spin configurations, as in our prior work [7–9]. Detector acceptance, trigger efficiency, and several other effects cancel in this ratio. Segmented beam-beam counters (BBCs) [21], symmetrically located on either side of the STAR interaction point and covering the pseudorapidity range  $3.4 < |\eta| < 5.0$ , measured the helicity-dependent relative luminosities and served as local polarimeters.

The STAR subsystems used to measure jets are the time projection chamber (TPC) and the barrel (BEMC) and endcap (EEMC) electromagnetic calorimeters [17]. The TPC provides tracking for charged particles in the 0.5 T solenoidal magnetic field with acceptance of  $|\eta| < 1.3$  and  $2\pi$  in the azimuthal angle  $\phi$ . The BEMC and EEMC cover a fiducial area of  $-1.0 < \eta < 2.0$  and  $0 < \phi < 2\pi$ , and they provide triggering and detection of photons and electrons.

Events were recorded if they satisfied the jet patch (JP) trigger condition in the BEMC or EEMC. The JP trigger required a  $\Delta\eta \times \Delta\phi = 1 \times 1$  patch of towers to exceed a transverse energy threshold of 5.4 (JP1, prescaled) or 7.3 (JP2) GeV, or two adjacent patches to each exceed 3.5 GeV (AJP). The addition of the AJP condition, combined with a reconfiguration of the jet patches so that they overlapped in  $\eta$ , resulted in a 37% increase in jet acceptance compared to previous data [9]. Upgrades in the data acquisition system allowed STAR to record events at much higher rates as well.

The analysis procedures were similar to those in Ref. [9] except where noted below. The inputs to the jet finder were the charged particle momenta measured by the TPC and the neutral energy depositions observed by the calorimeter towers. Jets were reconstructed using the anti- $k_T$  algorithm

[22], as implemented in the FastJet package [23], with a resolution parameter  $R = 0.6$ . This is a change from the midpoint cone algorithm [24] that was used in previous STAR inclusive jet analyses [7–9]. Anti- $k_T$  jets are less susceptible to diffuse soft backgrounds from underlying event and pileup contributions, which provides a significant reduction in the trigger bias described below.

Most frequently, charged hadrons deposit energy equivalent to a minimum ionizing particle (MIP) in the calorimeter towers. Because the TPC reconstructs the momentum of all charged particles, the inclusion of tower energy from charged hadrons results in an overestimation of the jet momentum. Fluctuations in the deposited tower energy when charged hadrons interact with calorimeter materials further distort the jet momentum and degrade the jet momentum resolution. In previous STAR jet analyses [7–9], this hadronic energy was accounted for by subtracting energy corresponding to a MIP from the energy deposited in any BEMC or EEMC tower with a charged track passing through it, and then using simulations to estimate the residual correction. In this analysis, the  $E_T$  of the matched tower was adjusted by subtracting either  $p_{TC}$  of the charged track or  $E_T$ , whichever was less. This procedure reduces the residual jet momentum corrections. It also reduces the sensitivity to fluctuations in the hadronic energy deposition, resulting in an improved jet momentum resolution of  $\approx 18\%$  compared to  $\approx 23\%$  in previous analyses. The bottom panel of Fig. 1 demonstrates that this new “ $p_T$  subtraction” scheme leads to an average for the neutral energy fraction (NEF) of the jet energy that is close to the value of about 1/3 expected from isospin considerations.

In this analysis, jets were required to have transverse momentum  $p_T > 5$  GeV/ $c$  and  $|\eta| < 1.0$ . Noncollision backgrounds such as beam-gas interactions and cosmic rays, observed as neutral energy deposits in the BEMC and EEMC, were minimized by requiring the NEF to be less than 0.94. Only jets that pointed to a triggered jet patch were considered. The top panel in Fig. 1 demonstrates the effect of the calorimeter trigger on the jet NEF. The trigger requirement skews the sample to larger neutral energies, especially for jets reconstructed near the trigger threshold. The lower panel shows that this bias is minimized by the  $p_T$  subtraction when the jet  $p_T$  is well above threshold.

Simulated events are used to calculate the jet momentum corrections and to estimate the systematic uncertainties. This analysis utilized simulated QCD events generated using the Perugia 0 tune [25] in PYTHIA 6.425 [26]. The PYTHIA events were processed through the STAR detector response package based on GEANT 3 [27] and then embedded into randomly triggered events. As a result, the TPC tracks and calorimeter hits reconstructed in the simulation sample incorporate the same beam background and pileup contributions as the data sample, providing excellent agreement between the data and simulation as shown in Fig. 1.

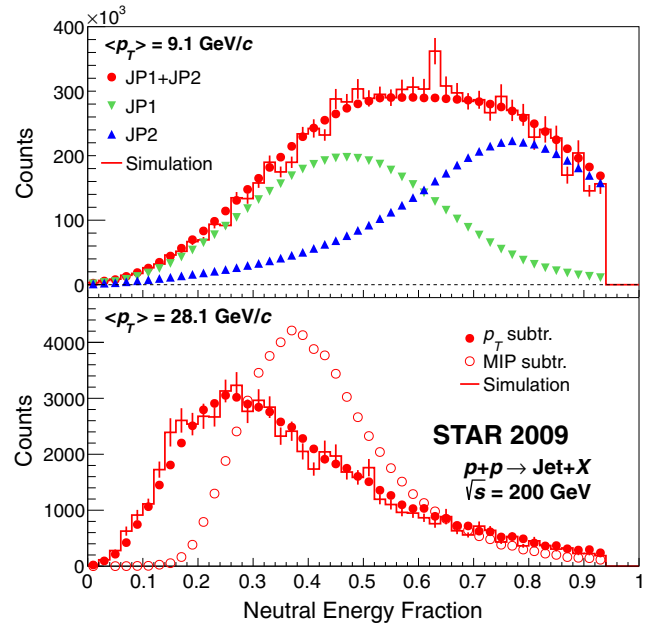


FIG. 1 (color online). Jet neutral energy fraction (NEF) comparing data (the solid points) with simulations (the histograms), where both are calculated with  $p_T$  subtraction. The upper panel shows jets with  $8.4 < p_T < 9.9$  GeV/ $c$ , demonstrating the bias in NEF when jet  $p_T$  is near the trigger threshold. The lower panel shows jets with  $26.8 < p_T < 31.6$  GeV/ $c$ , demonstrating that an apparent bias persists well above threshold when using MIP subtraction (the open circles). The error bars show the simulation statistics. Those for the data are smaller than the points.

The jet  $p_T$  reconstructed at the detector level can be corrected to either the particle or the parton level. Detector jets, which are formed from charged tracks and calorimeter towers, provide contact between the data and simulation. Particle jets are formed from the stable final-state particles produced in a collision. Parton jets are formed from the hard-scattered partons produced in the collision, including those from initial- and final-state radiation, but not those from the underlying event or beam remnants. Previous STAR analyses [7–9] corrected the data back to the particle level. Here, we correct the data to the parton jet level because parton jets provide a better representation of the jets in a NLO pQCD calculation. The anti- $k_T$  algorithm with  $R = 0.6$  was used to reconstruct parton jets for the simulated PYTHIA events described above. Simulated detector jets were matched to the parton jet closest in  $\eta - \phi$  space and within  $\sqrt{\Delta\eta^2 + \Delta\phi^2} \leq 0.5$ . Association probabilities ranged from 76% at the lowest jet  $p_T$  to  $> 98\%$  for  $p_T > 9.9$  GeV/ $c$ . Asymmetry values are given at the average parton jet  $p_T$  for each detector jet  $p_T$  bin.

The asymmetry  $A_{LL}$  was evaluated according to

$$A_{LL} = \frac{\sum (P_B P_Y)(N^{++} - rN^{+-})}{\sum (P_B P_Y)^2 (N^{++} + rN^{+-})}, \quad (2)$$

in which  $P_{B,Y}$  are the measured beam polarizations,  $N^{++}$  and  $N^{+-}$  denote the inclusive jet yields for equal and opposite proton beam helicity configurations, and  $r$  is the relative luminosity. Each sum is over individual runs that were 10 to 60 minutes long, a period much shorter than typical time variations in critical quantities such as  $P_{B,Y}$  and  $r$ . Values of  $r$  were measured run by run, and they range from 0.8 to 1.2.

The STAR trigger biases the data sample by altering the subprocess fractional contributions ( $gg$  vs  $qg$  vs  $qq$ ). At low  $p_T$ , the JP efficiency for quark jets is approximately 25% larger than for gluon jets. For  $p_T > 20$  GeV/ $c$ , the differences are negligible. Similarly, detector and trigger resolutions may smear and distort the measured  $A_{LL}$  values. The size of these effects depends on the value and the shape of the polarized gluon distribution as a function of Bjorken  $x$ . The  $A_{LL}$  values for detector jets were corrected for trigger and reconstruction bias effects by using the simulation to compare the observed asymmetries at the detector and parton jet levels. The PYTHIA event generator does not have options to simulate polarization effects in proton-proton collisions, but asymmetries can be constructed by using the kinematics of the hard interaction to access polarized and unpolarized parton distribution functions (PDFs) and calculate the expected asymmetry on an event-by-event basis. In this way, the trigger and reconstruction biases were calculated for a range of polarized PDFs that bracket the measured  $A_{LL}$  values. The average of the minimum and maximum  $A_{LL}^{\text{parton}} - A_{LL}^{\text{detector}}$  values for each jet  $p_T$  bin was used to correct the measured  $A_{LL}$  by amounts ranging from 0.0002 at low  $p_T$  to 0.0011 at high  $p_T$ , and half the difference was assigned as a (correlated) systematic uncertainty.

Figure 2 shows the inclusive jet  $A_{LL}$  plotted as a function of parton jet  $p_T$  for two  $\eta$  bins. The vertical size of the shaded uncertainty bands on the  $A_{LL}$  points in Fig. 2 reflects the quadrature sum of the systematic uncertainties due to corrections for the trigger and reconstruction bias ( $2\text{--}55 \times 10^{-4}$ ) and asymmetries associated with the residual transverse polarizations of the beams ( $3\text{--}26 \times 10^{-4}$ ). The trigger and reconstruction bias contributions are dominated by the statistics of the simulation sample. The residual transverse polarization contributions are dominated by the statistical uncertainties in the measurement of the relevant transverse double-spin asymmetry ( $A_{\Sigma}$ ) [9]. Both of these uncertainties are primarily point-to-point fluctuating. Contributions to  $A_{LL}$  from noncollision backgrounds were estimated to be less than 2% of the statistical uncertainty on  $A_{LL}$  for all jet  $p_T$  bins and deemed negligible. Likewise, uncertainties associated with the possible dependence of the underlying event on the configuration of the beam spins were neglected. The relative luminosity uncertainty ( $\pm 5 \times 10^{-4}$ ), which is common to all of the points, is shown by the gray bands on the horizontal axes. It was estimated by comparing the relative luminosities calculated with the BBCs and

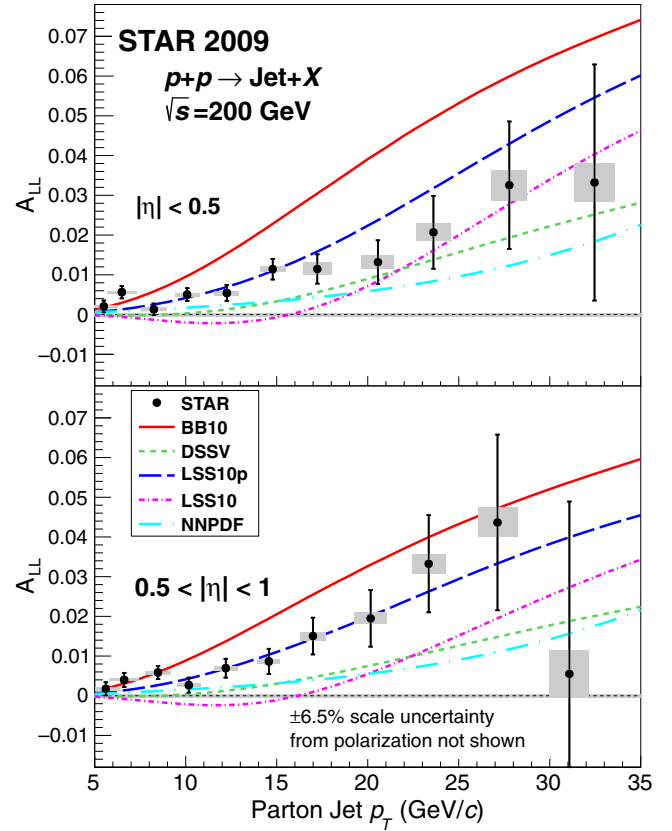


FIG. 2 (color online). Midrapidity ( $|\eta| < 0.5$ , upper panel) and forward rapidity ( $0.5 < |\eta| < 1$ , lower panel) inclusive jet  $A_{LL}$  vs parton jet  $p_T$ , compared to predictions from several NLO global analyses. The error bars are statistical. The gray boxes show the size of the systematic uncertainties.

zero-degree calorimeters [17], and from inspection of a number of asymmetries expected to yield null results. The horizontal size of the shaded error bands reflects the systematic uncertainty on the corrected jet  $p_T$ . This includes calorimeter tower gain and efficiency and TPC tracking efficiency and momentum resolution effects. An additional uncertainty has been added in quadrature to account for the difference between the PYTHIA parton jet and NLO pQCD jet cross sections. The PYTHIA vs NLO pQCD difference dominates for most bins, making the parton jet  $p_T$  uncertainties highly correlated. The values for  $A_{LL}$  by  $p_T$  bin, their uncertainties, and correlations are given in the Supplemental Material [28] together with particle and parton jet  $p_T$ .

Longitudinal single-spin asymmetries,  $A_L$ , measure parity-violating effects arising from weak interactions, and hence are expected to be very small compared to  $A_{LL}$  at  $\sqrt{s} = 200$  GeV.  $A_L$  was measured and found to be consistent with zero for each beam, as expected for the present data statistics.

The theoretical curves in Fig. 2 illustrate the  $A_{LL}$  expected for the polarized PDFs associated with the corresponding global analyses. These predictions were made by inserting the polarized PDFs from BB [4],

DSSV [2,3], LSS [5], and NNPDF [6] into the NLO jet production code of Mukherjee and Vogelsang [16]. Theoretical uncertainty bands for  $A_{LL}$  were also calculated, but they are omitted from the figure for clarity. The BB10 and NNPDF polarized PDFs are based only on inclusive DIS data, while LSS includes both inclusive and semi-inclusive DIS (SIDIS) data sets. LSS provides two distinct solutions for the polarized gluon density of nearly equal quality. The LSS10 gluon density has a node at  $x \approx 0.2$ , and the LSS10p gluon is positive definite at the input scale  $Q_0^2 = 2.5 \text{ GeV}^2$ . DSSV is the only fit that incorporates DIS, SIDIS, and previous RHIC  $pp$  data.

LSS10p provides a good description of these STAR jet data. The STAR results lie above the predictions of DSSV and NNPDF and below the predictions of BB10. However, the measurements fall within the combined data and model uncertainties for these three cases. In contrast, the STAR jet asymmetries are systematically above the predictions of LSS10 and fall outside the LSS10 uncertainty band for  $p_T < 15 \text{ GeV}/c$ . The quantum statistical parton distribution approach [29,30] now incorporates a positive gluon polarization and obtains very reasonable agreement with our data.

The DSSV group has performed a new global analysis [31] including the STAR jet  $A_{LL}$  results reported in this Letter. They find that the integral of  $\Delta g(x, Q^2 = 10 \text{ GeV}^2)$  over the range  $x > 0.05$  is  $0.20_{-0.07}^{+0.06}$  at 90% C.L. DSSV indicates that the STAR jet data lead to the positive gluon polarization in the RHIC kinematic range.

The NNPDF group follows a conceptually different approach and has developed a reweighting method [32,33] to include new experimental data into an existing PDF set without the need to repeat the entire fitting process. The method involves calculating weighted averages over previously equivalent PDF sets, with the weight for each set derived from the  $\chi^2$  probability for the set to describe the new data. In their recent work [34], NNPDF finds that the integral of  $\Delta g(x, Q^2 = 10 \text{ GeV}^2)$  over the range  $0.05 < x < 0.20$  is  $0.17 \pm 0.06$ . This is to be compared with the threefold less precise value of  $0.05 \pm 0.15$  prior to the inclusion of the present STAR jet data. The value over the range  $x > 0.05$  is  $0.23 \pm 0.06$  [34,35].

The recently published DSSV and NNPDF results are consistent. The functional form of the polarized parton distribution functions assumed by DSSV is less flexible than that assumed by NNPDF, and DSSV includes substantially more data in their fit. In both analyses, the inclusion of the STAR jet data results in a substantial reduction in the uncertainty for the gluon polarization in the region  $x > 0.05$  and indicates a preference for the gluon helicity contribution to be positive in the RHIC kinematic range.

In summary, we report a new measurement of the inclusive jet longitudinal double-spin asymmetry  $A_{LL}$  in polarized  $pp$  collisions at  $\sqrt{s} = 200 \text{ GeV}$ . The results are consistent with predictions from several recent NLO

polarized parton distribution fits. When included in updated global analyses, they provide evidence at the  $3\sigma$  level for positive gluon polarization in the region  $x > 0.05$ .

We would like to thank J. Blümlein, H. Böttcher, E. Leader, E. Nocera, D. B. Stamenov, M. Stratmann, and W. Vogelsang for information regarding their respective polarized PDF sets and their uncertainties. We thank the RHIC Operations Group and RCF at BNL, the NERSC Center at LBNL, the KISTI Center in Korea, and the Open Science Grid consortium for providing resources and support. This work was supported in part by the Offices of NP and HEP within the U.S. DOE Office of Science, the U.S. NSF, CNRS/IN2P3, FAPESP CNPq of Brazil, the Ministry of Education and Science of the Russian Federation, NNSFC, CAS, MoST and MoE of China, the Korean Research Foundation, GA and MSMT of the Czech Republic, FIAS of Germany, DAE, DST, and CSIR of India, the National Science Centre of Poland, the National Research Foundation (NRF-2012004024), the Ministry of Science, Education and Sports of the Republic of Croatia, and RosAtom of Russia.

- 
- [1] C. A. Aidala, S. D. Bass, D. Hasch, and G. K. Mallot, *Rev. Mod. Phys.* **85**, 655 (2013), and references therein.
  - [2] D. de Florian, R. Sassot, M. Stratmann, and W. Vogelsang, *Phys. Rev. Lett.* **101**, 072001 (2008).
  - [3] D. de Florian, R. Sassot, M. Stratmann, and W. Vogelsang, *Phys. Rev. D* **80**, 034030 (2009).
  - [4] J. Blümlein and H. Böttcher, *Nucl. Phys.* **B841**, 205 (2010).
  - [5] E. Leader, A. V. Sidorov, and D. B. Stamenov, *Phys. Rev. D* **82**, 114018 (2010).
  - [6] R. D. Ball *et al.* (NNPDF Collaboration), *Nucl. Phys.* **B874**, 36 (2013).
  - [7] B. I. Abelev *et al.* (STAR Collaboration), *Phys. Rev. Lett.* **97**, 252001 (2006).
  - [8] B. I. Abelev *et al.* (STAR Collaboration), *Phys. Rev. Lett.* **100**, 232003 (2008).
  - [9] L. Adamczyk *et al.* (STAR Collaboration), *Phys. Rev. D* **86**, 032006 (2012).
  - [10] S. S. Adler *et al.* (PHENIX Collaboration), *Phys. Rev. D* **73**, 091102 (2006).
  - [11] A. Adare *et al.* (PHENIX Collaboration), *Phys. Rev. D* **76**, 051106 (2007).
  - [12] A. Adare *et al.* (PHENIX Collaboration), *Phys. Rev. Lett.* **103**, 012003 (2009).
  - [13] A. Adare *et al.* (PHENIX Collaboration), *Phys. Rev. D* **79**, 012003 (2009).
  - [14] A. Adare *et al.* (PHENIX Collaboration), *Phys. Rev. D* **90**, 012007 (2014).
  - [15] B. Jäger, M. Stratmann, and W. Vogelsang, *Phys. Rev. D* **70**, 034010 (2004).
  - [16] A. Mukherjee and W. Vogelsang, *Phys. Rev. D* **86**, 094009 (2012).
  - [17] K. H. Ackermann *et al.* (STAR Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **499**, 624 (2003), and references therein.

- [18] O. Jinnouchi *et al.*, [arXiv:nucl-ex/0412053](#).
- [19] H. Okada *et al.*, [arXiv:hep-ex/0601001](#).
- [20] B. Schmidke *et al.*, RHIC Polarization for Runs 9–12, Brookhaven National Laboratory Collider-Accelerator Department Report No. C-A/AP/490, 2013, <http://public.bnl.gov/docs/cad/Pages/Home.aspx>.
- [21] J. Koryluk *et al.* (STAR Collaboration), [arXiv:hep-ex/0501072](#).
- [22] M. Cacciari, G. P. Salam, and G. Soyez, *J. High Energy Phys.* **04** (2008) 063.
- [23] M. Cacciari, G. P. Salam, and G. Soyez, *Eur. Phys. J. C* **72**, 1896 (2012).
- [24] G. C. Blazey *et al.*, [arXiv:hep-ex/0005012](#).
- [25] P. Z. Skands, [arXiv:0905.3418](#).
- [26] T. Sjostrand, S. Mrenna, and P. Z. Skands, *J. High Energy Phys.* **05** (2006) 026.
- [27] GEANT 3.21, CERN Program Library.
- [28] See Supplemental Material at <http://link.aps.org/supplemental/10.1103/PhysRevLett.115.092002> for tables of  $A_{LL}$  values by  $p_T$  bin, their uncertainties, and their correlations.
- [29] C. Bourrely, J. Soffer, and F. Buccella, *Eur. Phys. J. C* **23**, 487 (2002).
- [30] C. Bourrely and J. Soffer, *Phys. Lett. B* **740**, 168 (2015).
- [31] D. de Florian, R. Sassot, M. Stratmann, and W. Vogelsang, *Phys. Rev. Lett.* **113**, 012001 (2014).
- [32] R. D. Ball, V. Bertone, F. Cerutti, L. Del Debbio, S. Forte, A. Guffanti, J. I. Latorre, J. Rojo, and M. Ubiali (NNPDF Collaboration), *Nucl. Phys.* **B849**, 112 (2011); **B854**, 926 (E) (2012); **B855**, 927(E) (2012).
- [33] R. D. Ball, V. Bertone, F. Cerutti, L. Del Debbio, S. Forte, A. Guffanti, N. P. Hartland, J. I. Latorre, J. Rojo, and M. Ubiali (NNPDF Collaboration), *Nucl. Phys.* **B855**, 608 (2012).
- [34] E. R. Nocera *et al.* (NNPDF Collaboration), *Nucl. Phys.* **B887**, 276 (2014).
- [35] E. R. Nocera, [arXiv:1503.03518](#); (private communication).